

On the influence of ocean dynamics on gravity noise

Papp Gábor, Szűcs Eszter, Battha László, Benedek Judit

Geodetic and Geophysical Institute of the
Research Centre for Astronomy and Earth Sciences
Hungarian Academy of Sciences
Sopron - *Civitas Fidelissima*

Draft of the presentation

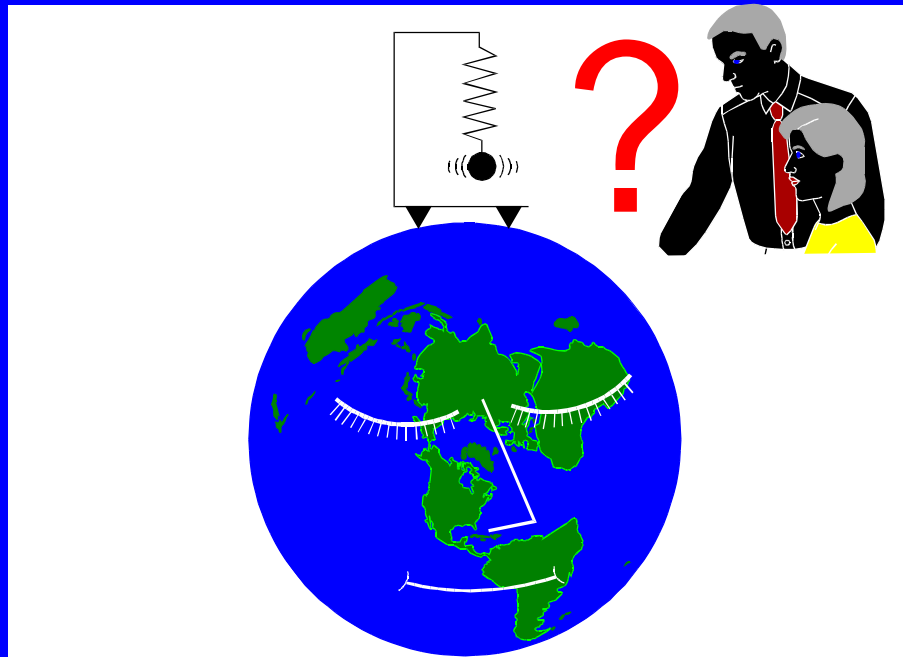
- "Eppur si muove..." - the lament of *homo gravimetricus*
- Non-gravitational effects in **g** observations
- Comparison of gravity tide and microseismic noise
- Level of the gravity noise observed in Sopronbánfalva Observatory (West Hungary) between 08.06.2010 és 04.01.2011
- Estimation of the sensitivity of gravimetric observations based on spectral analysis of co-located seismological records
- Identification of the sources of microseisms
- Spectral analysis of 1 Hz records (*preliminary results*)

Eppur si muove

- The lament of *homo gravimetricus*,

i.e.

Why does his instrument tremble when no earthquake can be felt?

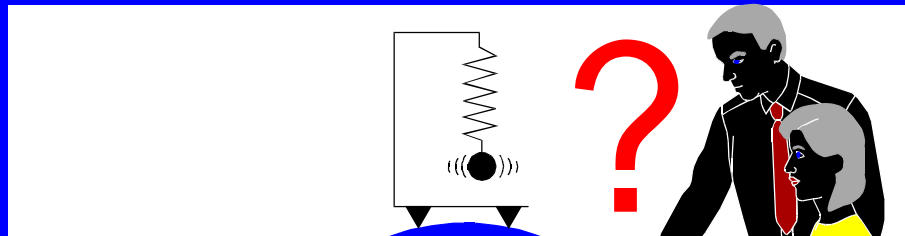


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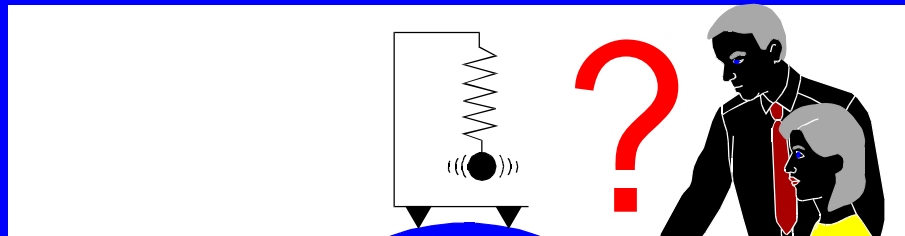
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- B) If $\neg A$ then there is microseismic activity
- C) If $\neg(A \wedge B)$ then there are road construction works close to the observatory

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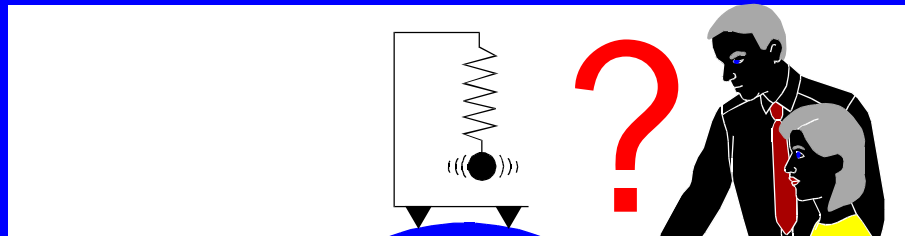
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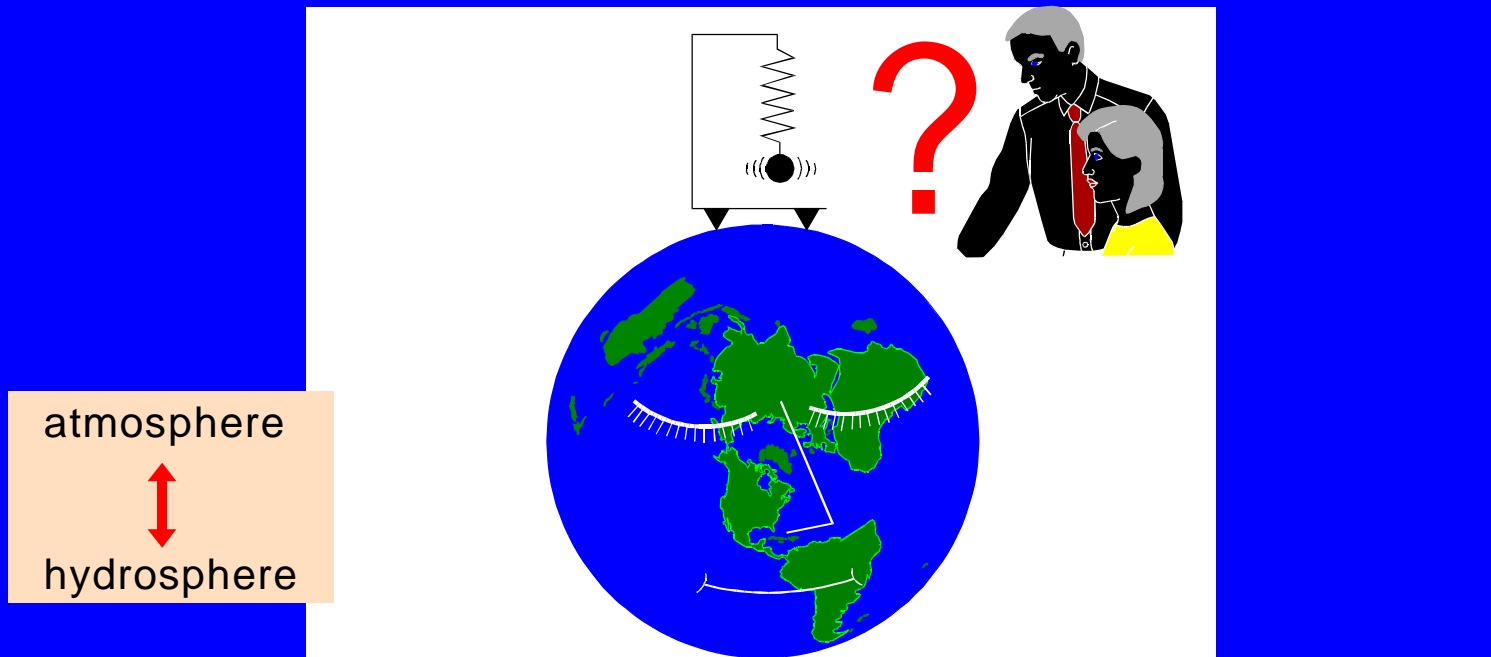
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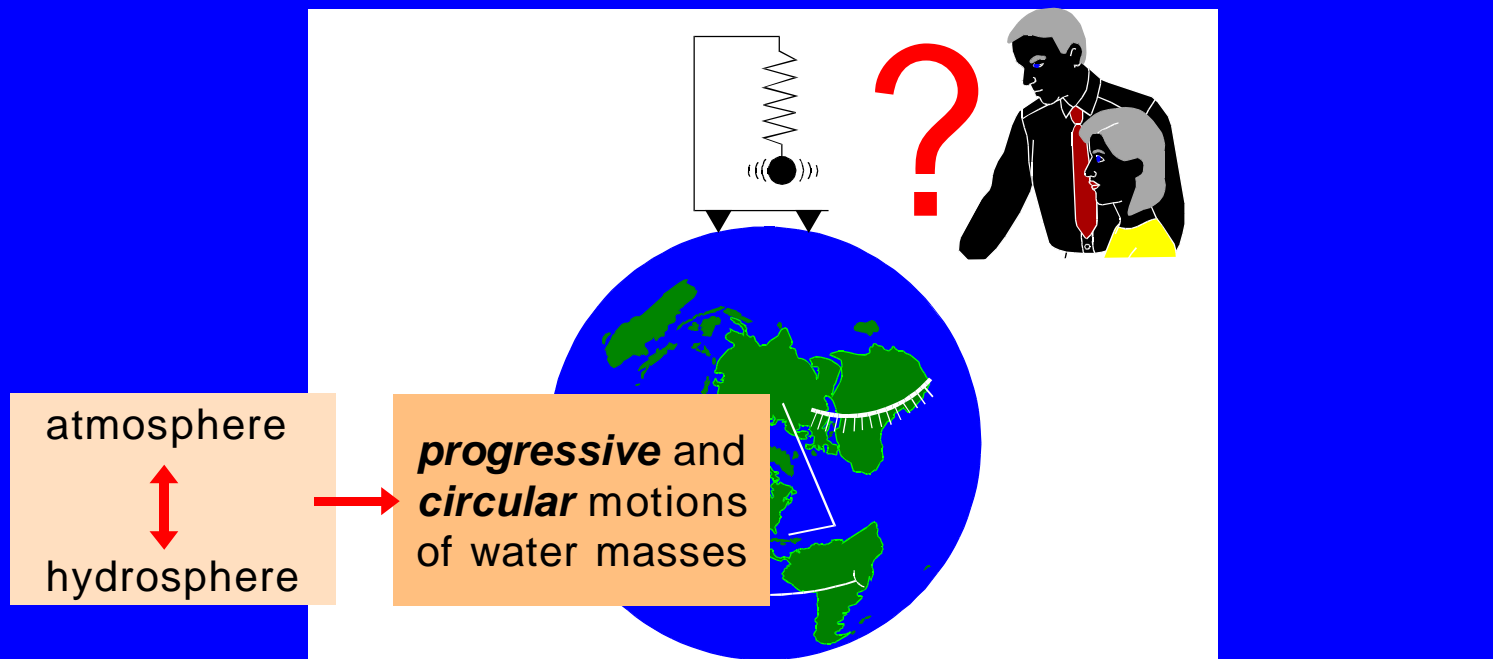


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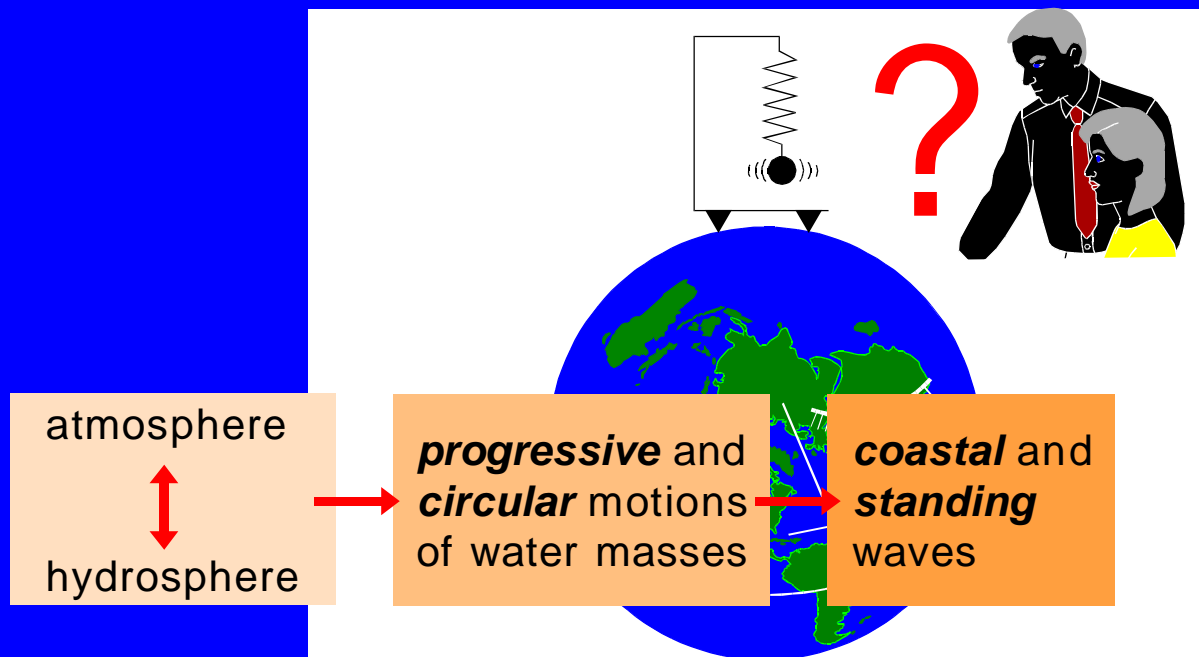


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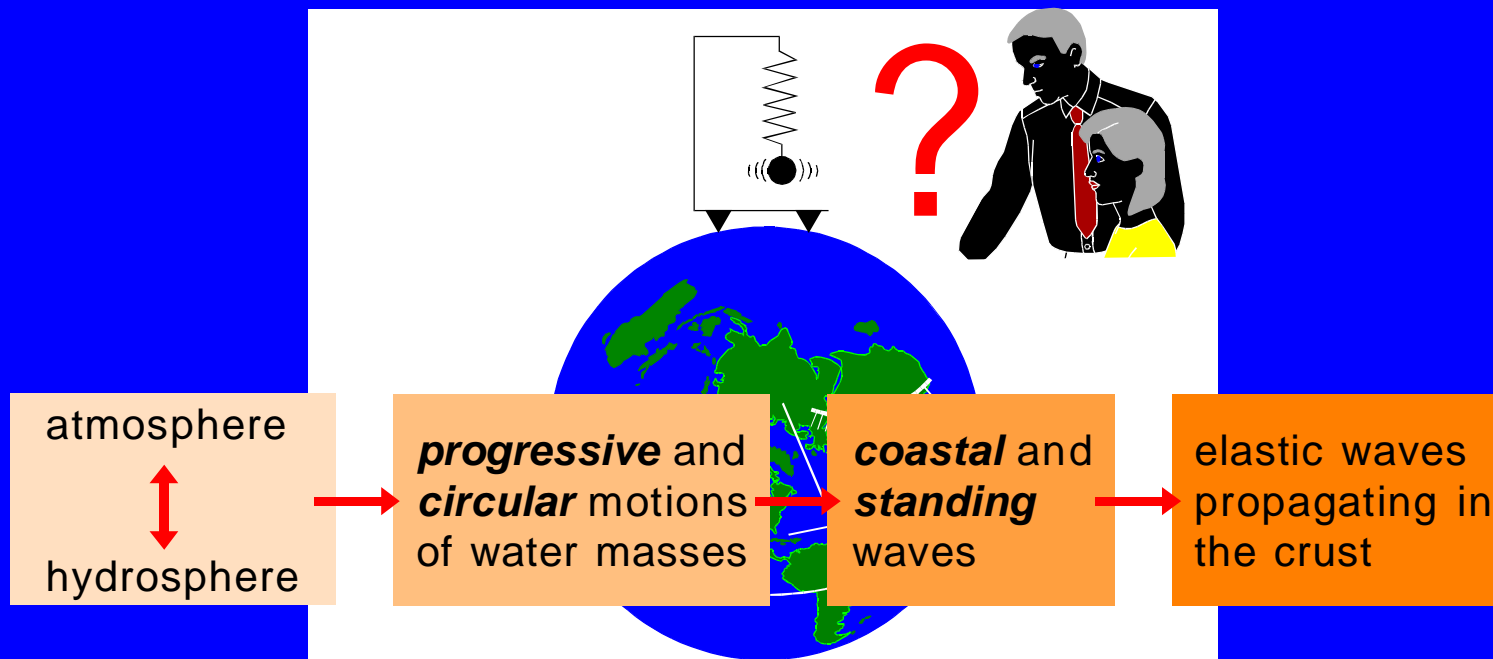


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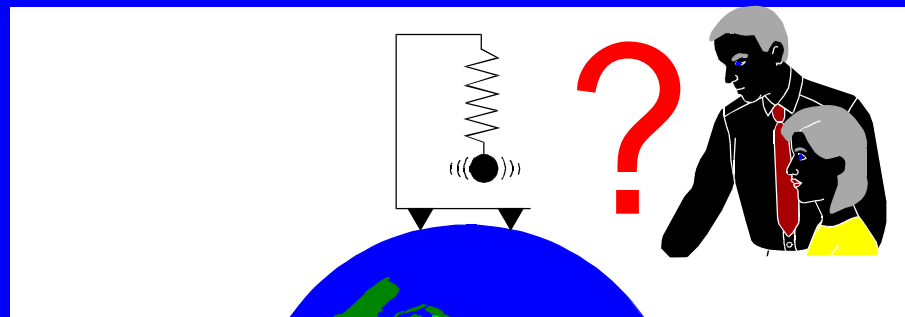


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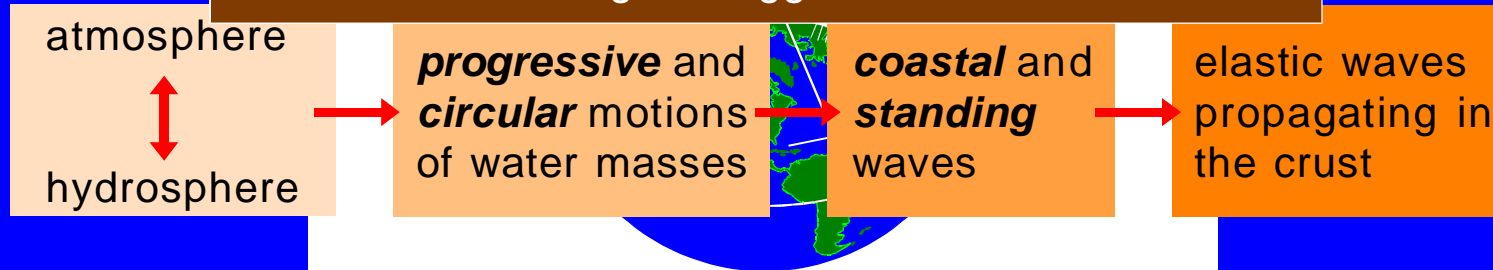
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Longuet-Higgins, 1950

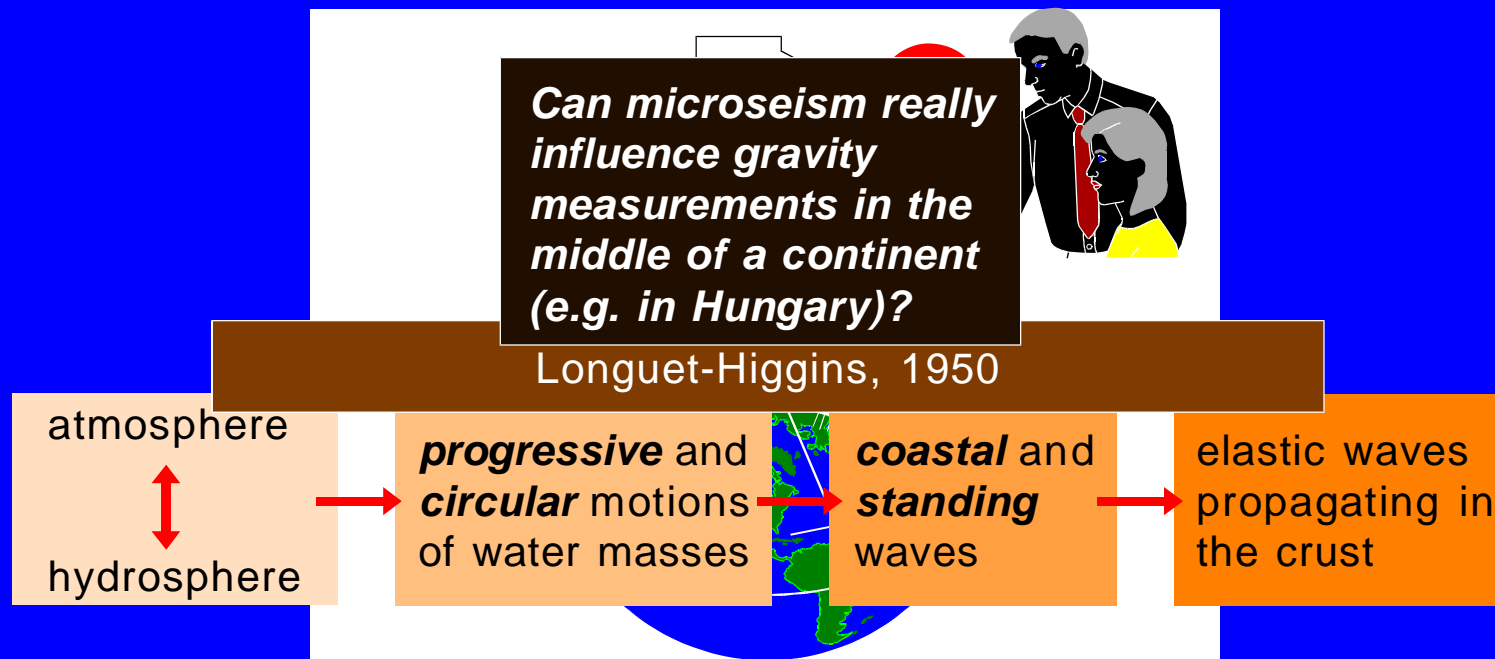


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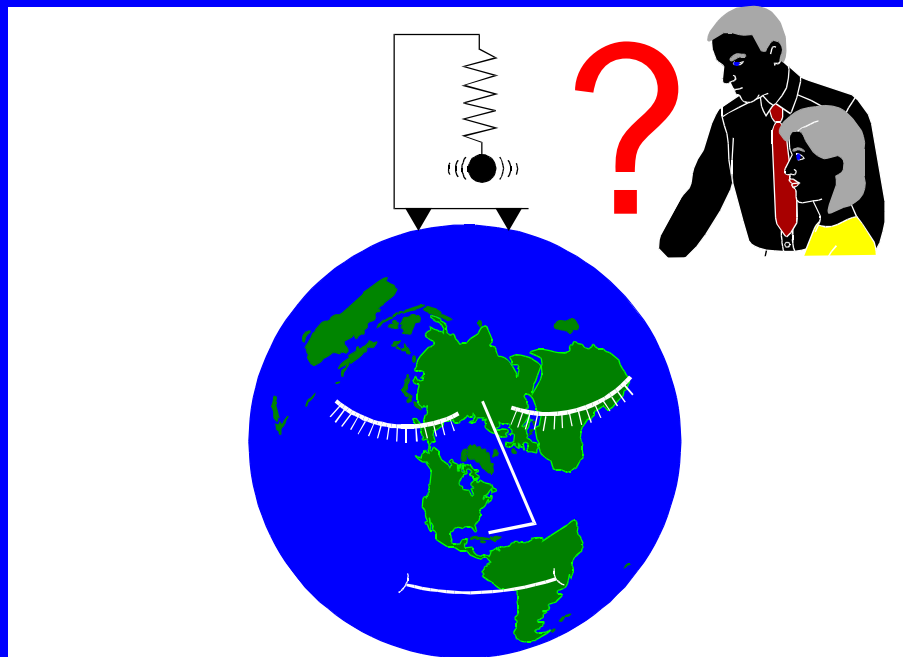


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The sum of time dependent effects (I) due to the change in motion of the observation point/instrument

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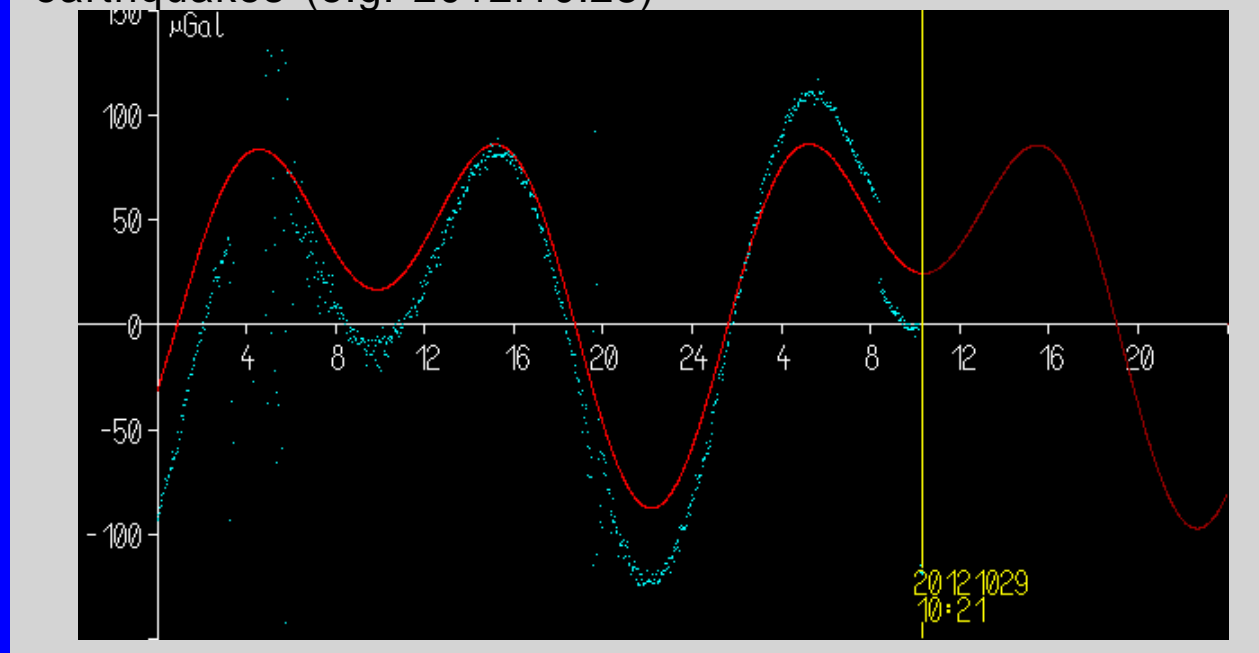
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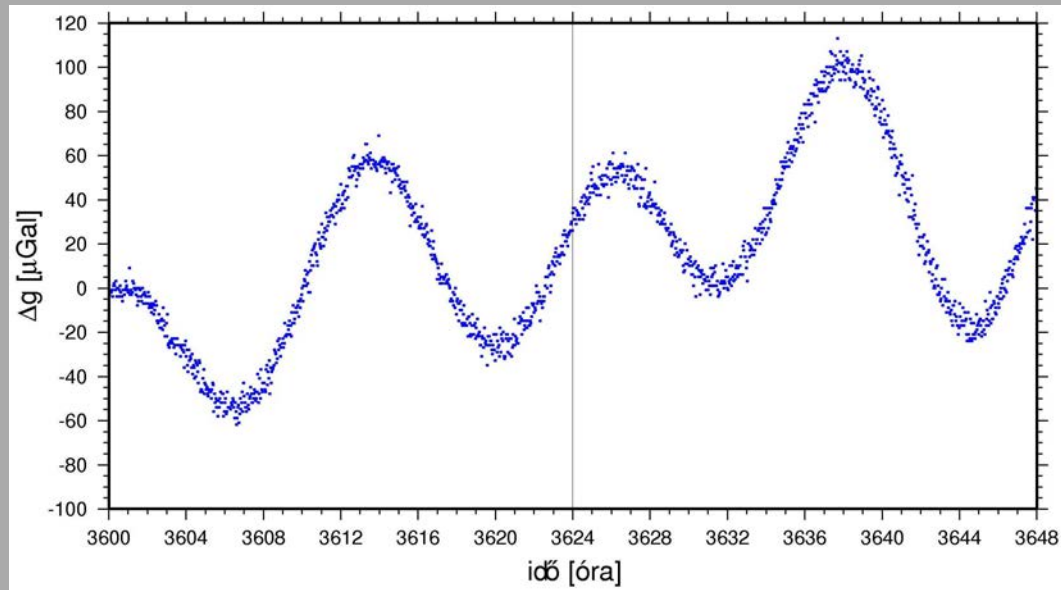


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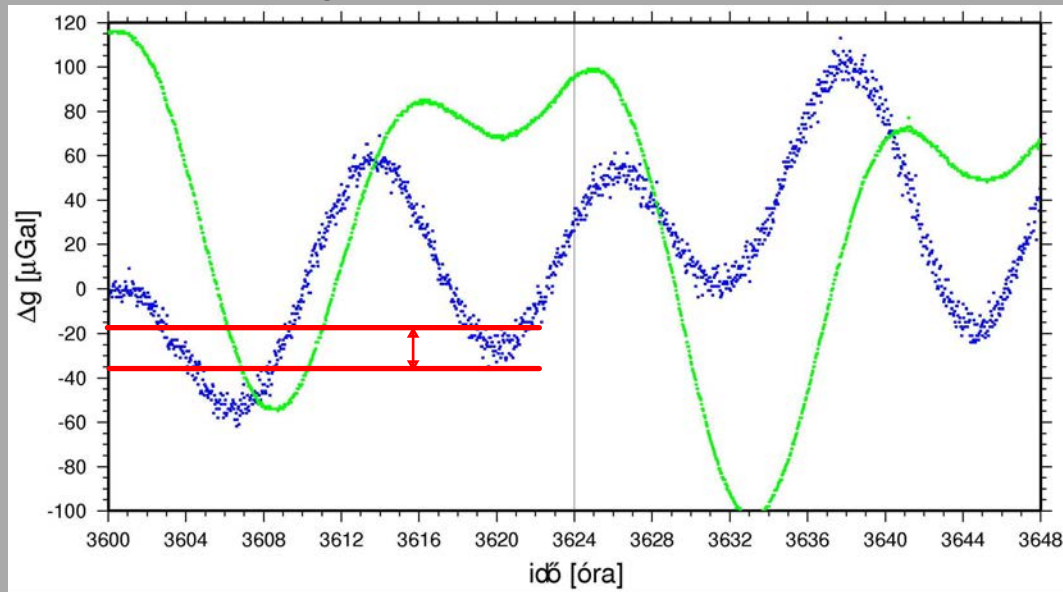


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Comparison of gravity tide to microseismic noise

- Spectral decomposition of gravity tide

The most dominant components of gravity tide at $\Phi=47.6^\circ$
(Baker, 1984)

component	description	T_i cycle time [h/day]	a_i amplitude [μGal]
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N_2	lunar elliptic	12.66	6.5
K_2	lunisolar decl.	11.97	4.3
O_1	diurnal lunar	25.82	30.9
K_1	lunisolar decl.	23.93	43.5
P_1	diurnal solar	24.07	14.4
M_f	lunar fortnightly	13.66 day	4.1
M_m	lunar monthly	27.55 day	2.1
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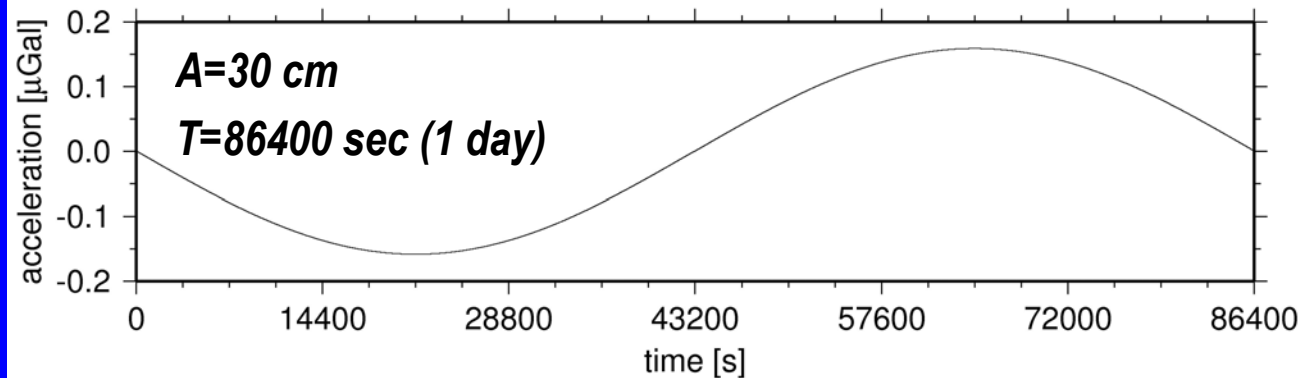
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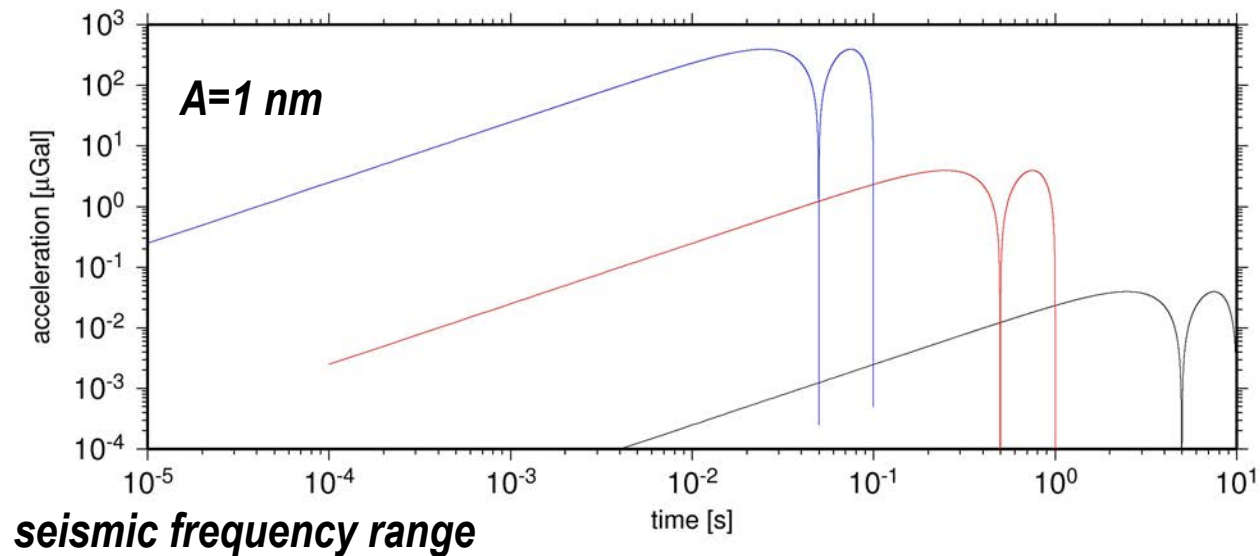
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daily scale factor

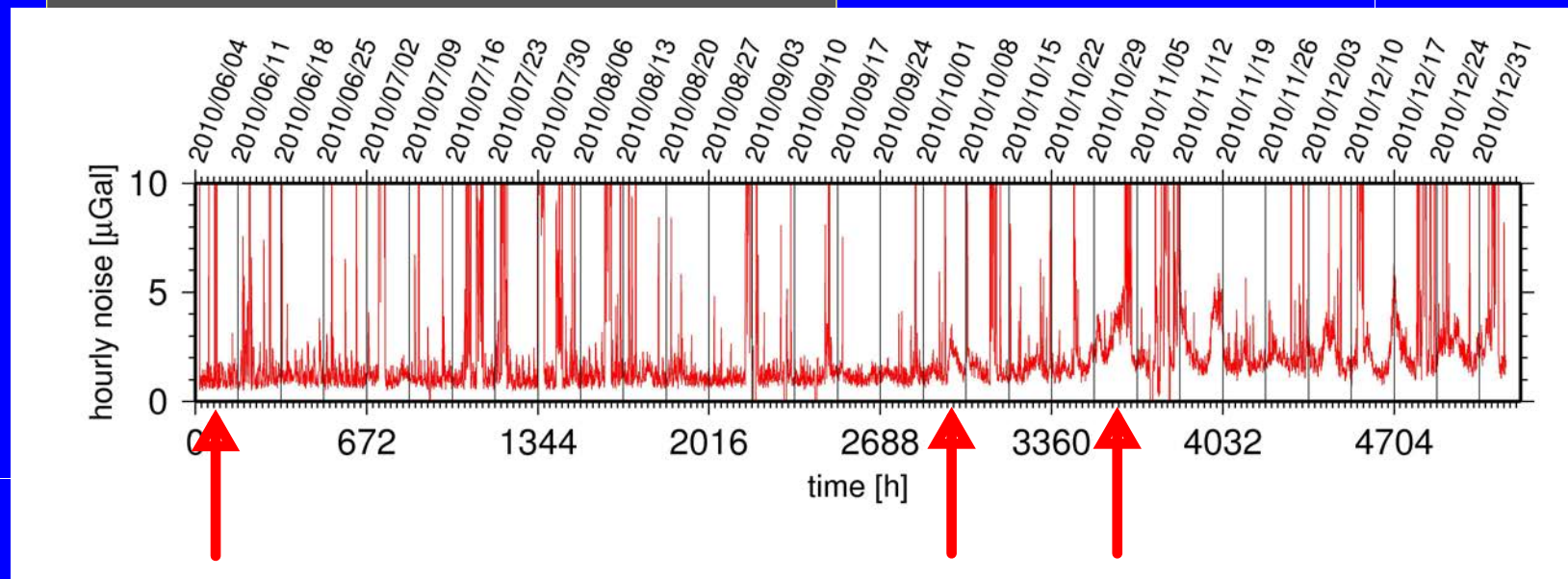
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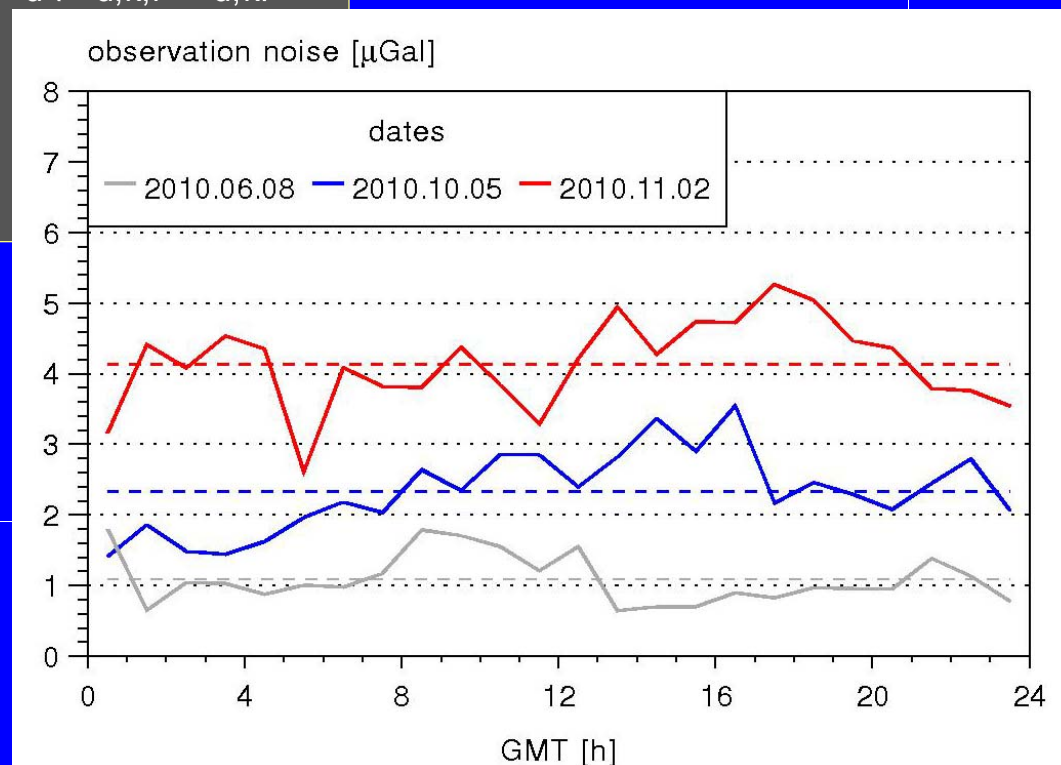
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Sensitivity estimation of **g** observations

Co-located measurements with SOP station of the Hungarian National Seismological Network
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 $\Delta t = t_{i+1} - t_i = 0.05 \text{ s (20 Hz)}$

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amplitude Fourier spectrum

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maximum inertial acceleration

$$a_{\max}(A, T) = \text{abs}(-4A\pi^2/T^2)$$

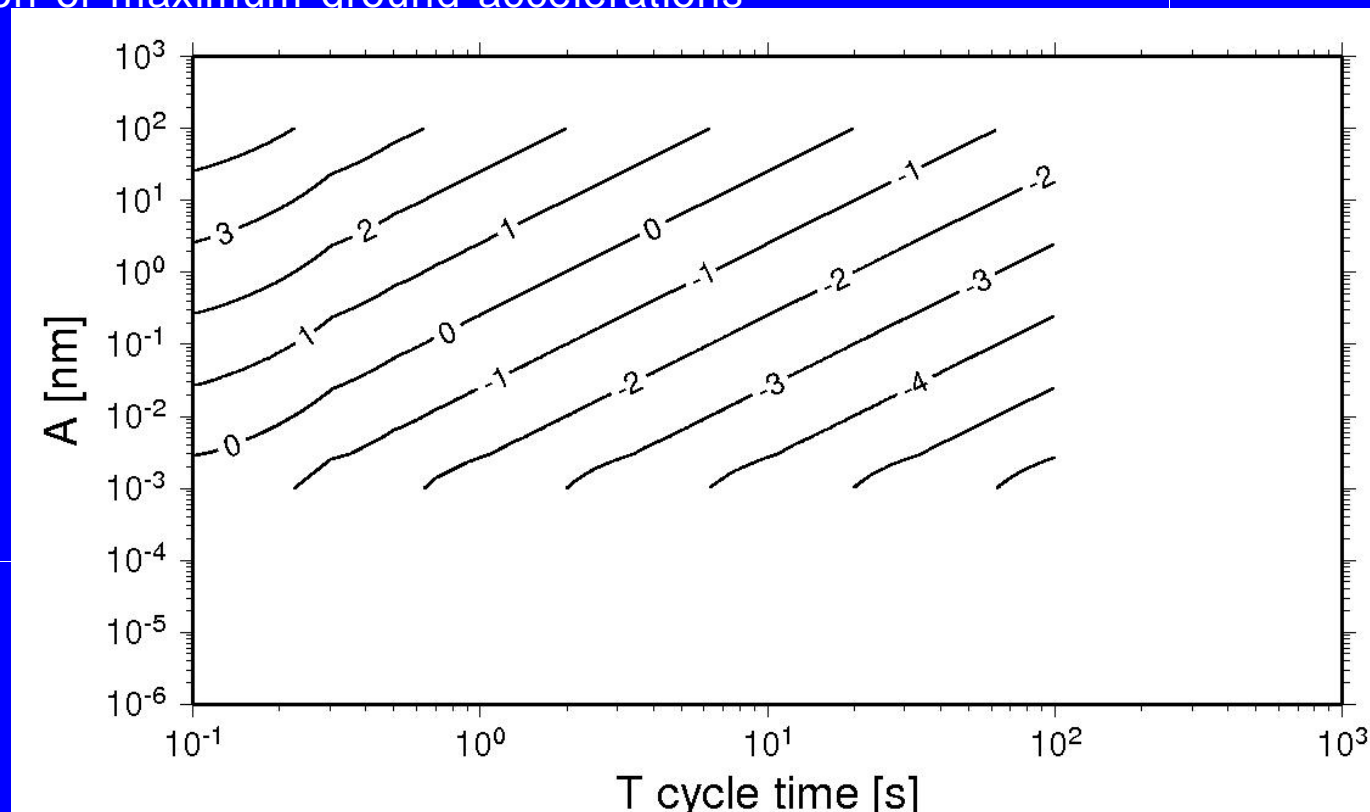
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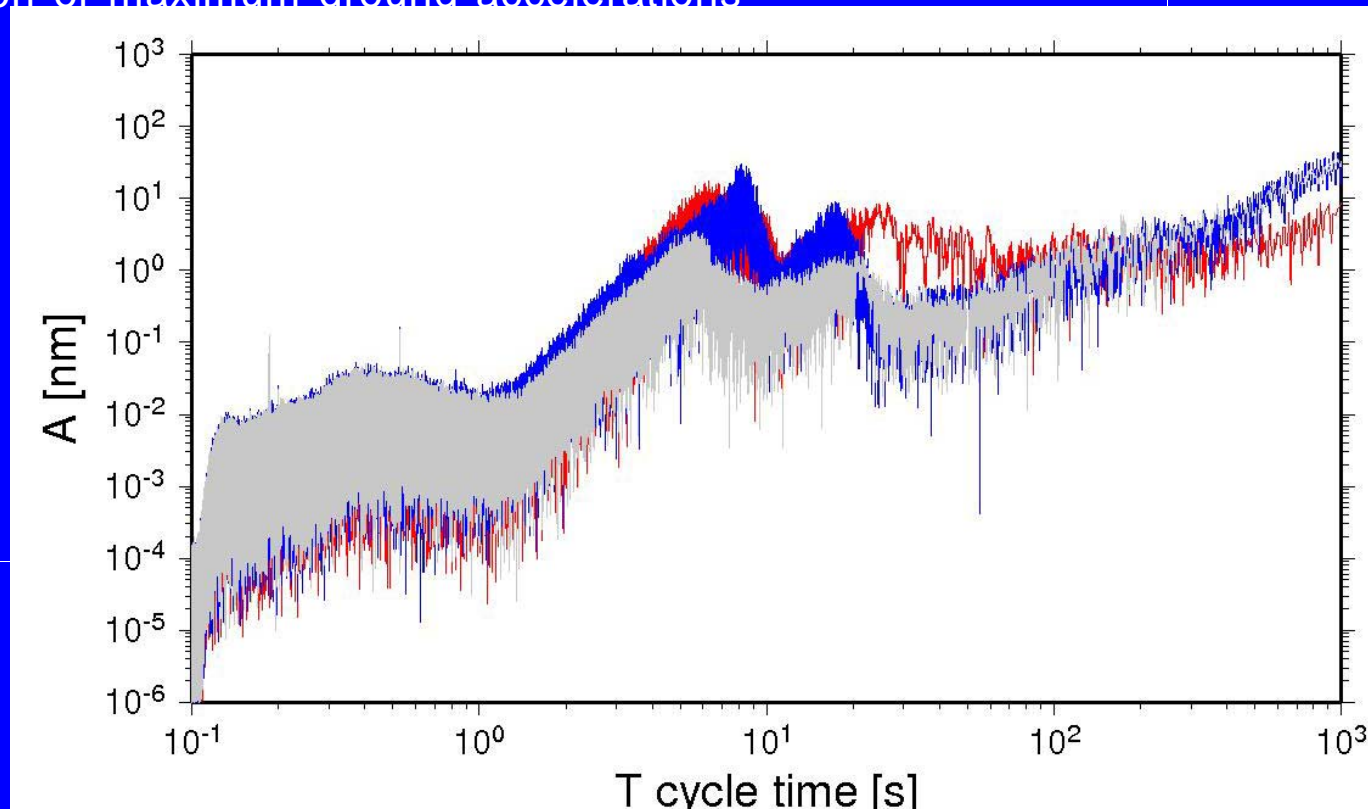
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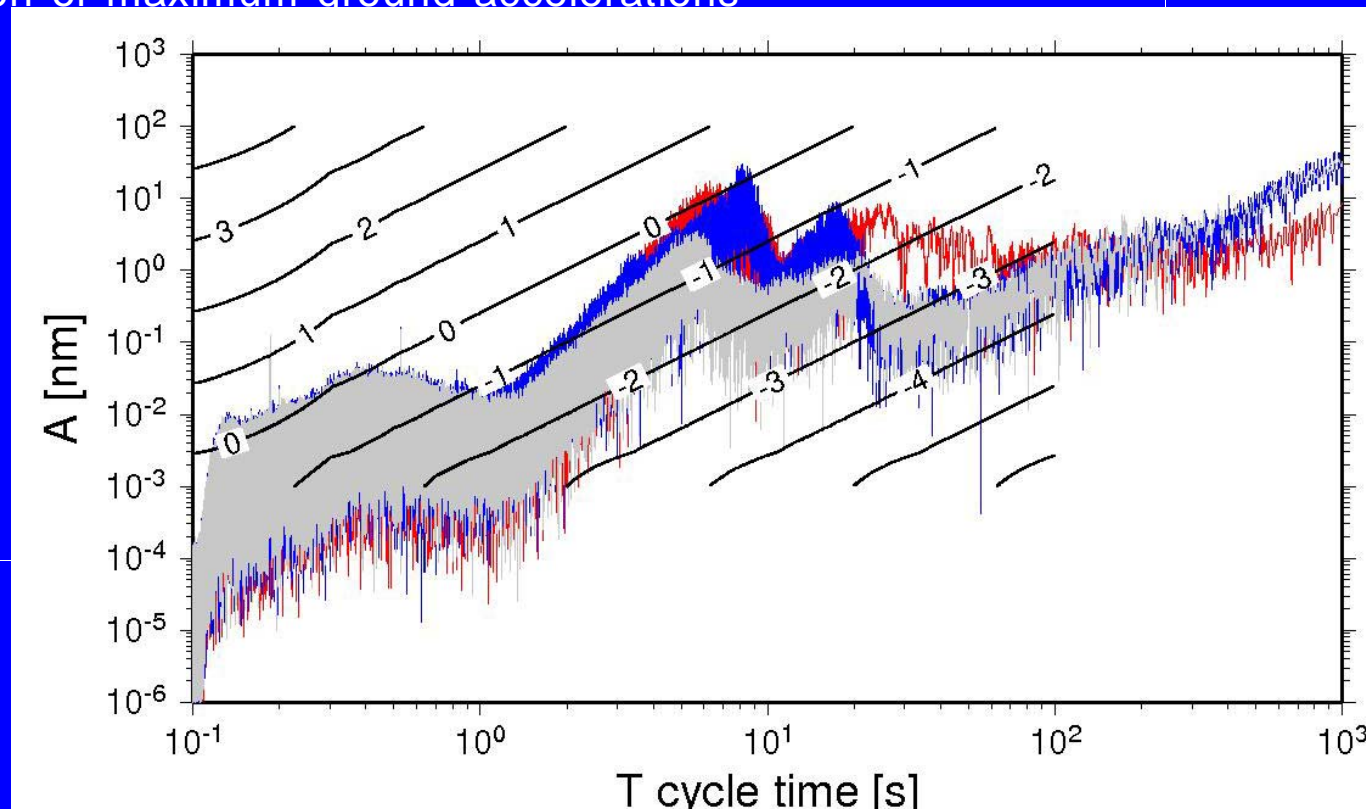
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Identification of the sources of microseisms

- The origin of microseismic "noise" (Longuet-Higgins, 1950)

Dynamical interactions between the hydrosphere and atmosphere

Surface waves generated by the swelling of ocean's surface

1) coastal waves
(primary source)
 $10 \text{ s} < T < 16 \text{ s}$

2) standing waves
(secondary source)
 $4 \text{ s} < T < 8 \text{ s}$

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- Monitoring of ocean weather (e.g. Oceanweather Inc. USA)

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weather maps (4 maps/day)
- H_{sw} - significant wave height
- wave direction

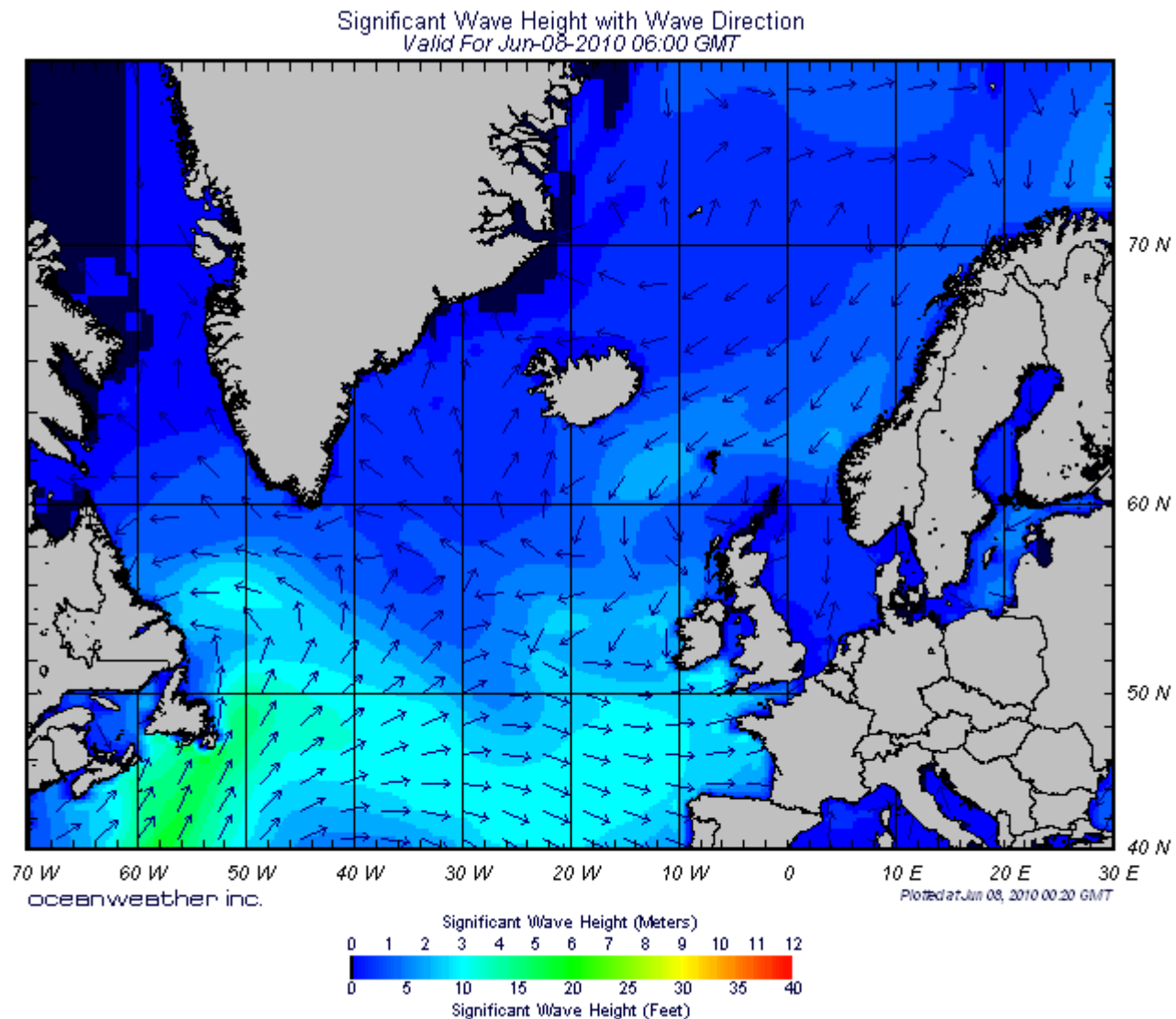
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- The origin
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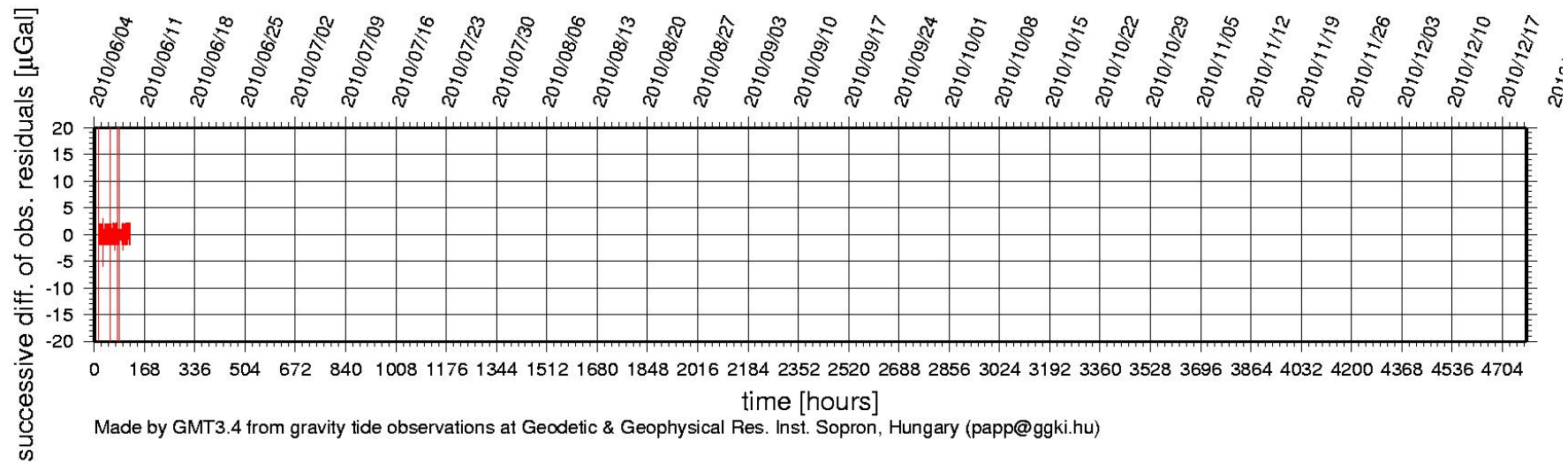
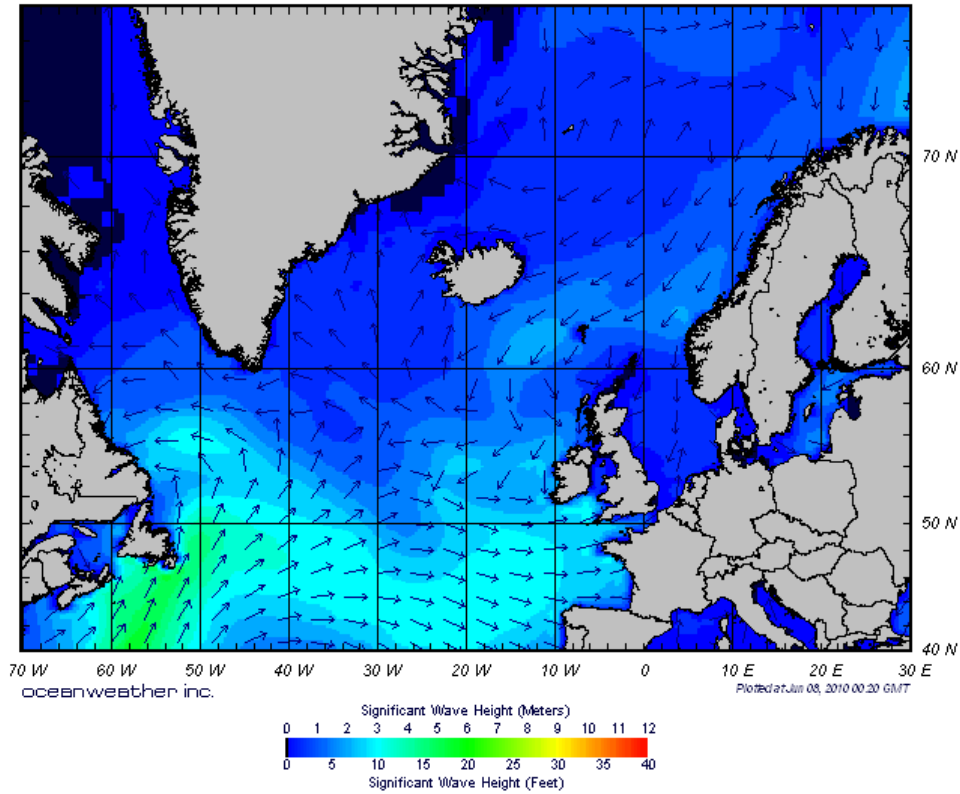
wea

- H

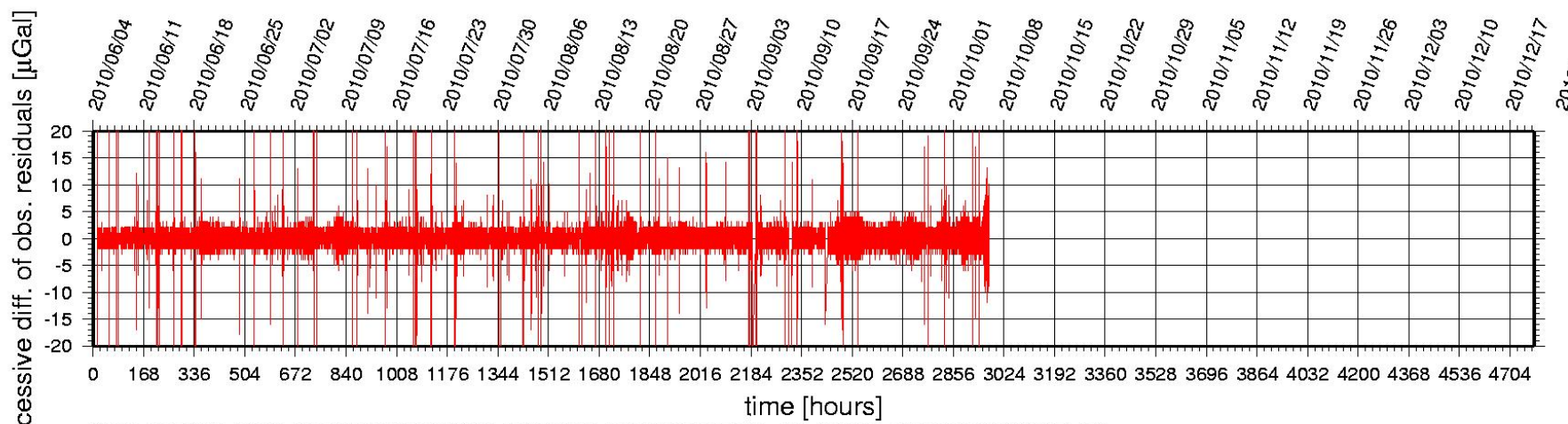
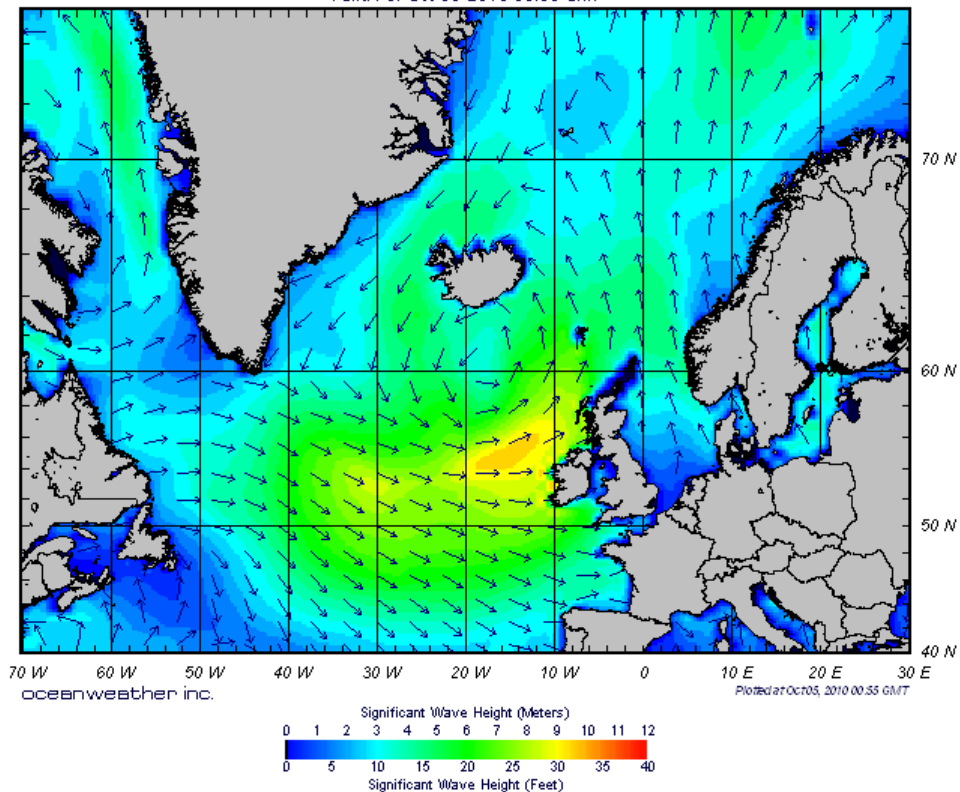
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Significant Wave Height with Wave Direction
Valid For Jun-08-2010 06:00 GMT

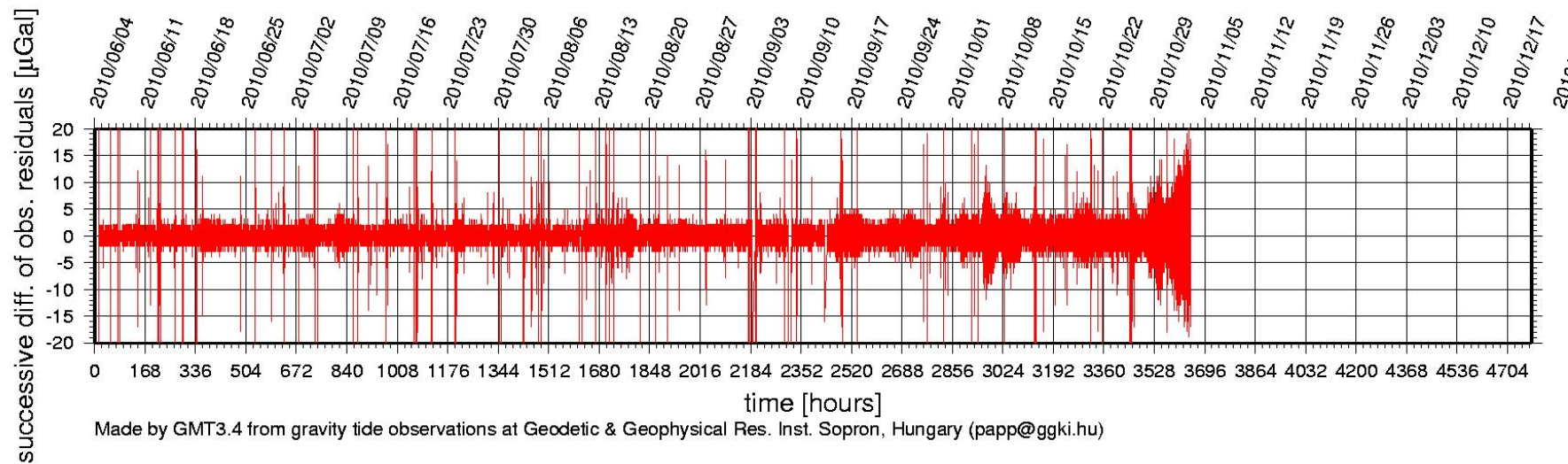
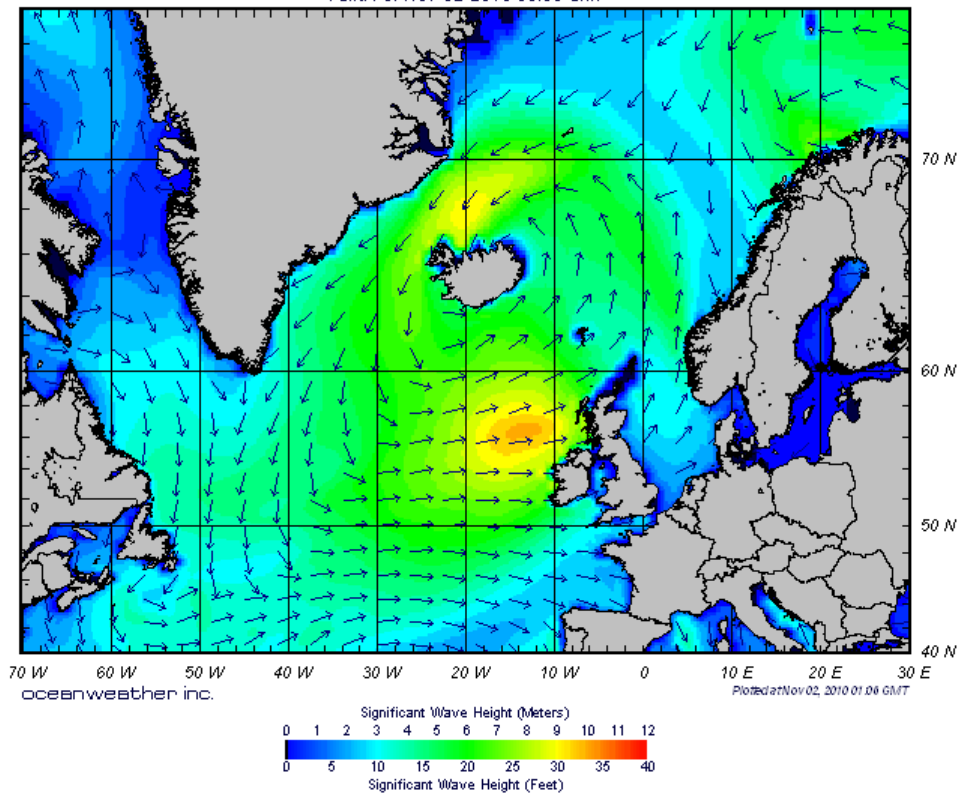


Significant Wave Height with Wave Direction
Valid For Oct-05-2010 06:00 GMT



Made by GMT3.4 from gravity tide observations at Geodetic & Geophysical Res. Inst. Sopron, Hungary (papp@ggki.hu)

Significant Wave Height with Wave Direction
Valid For Nov-02-2010 06:00 GMT

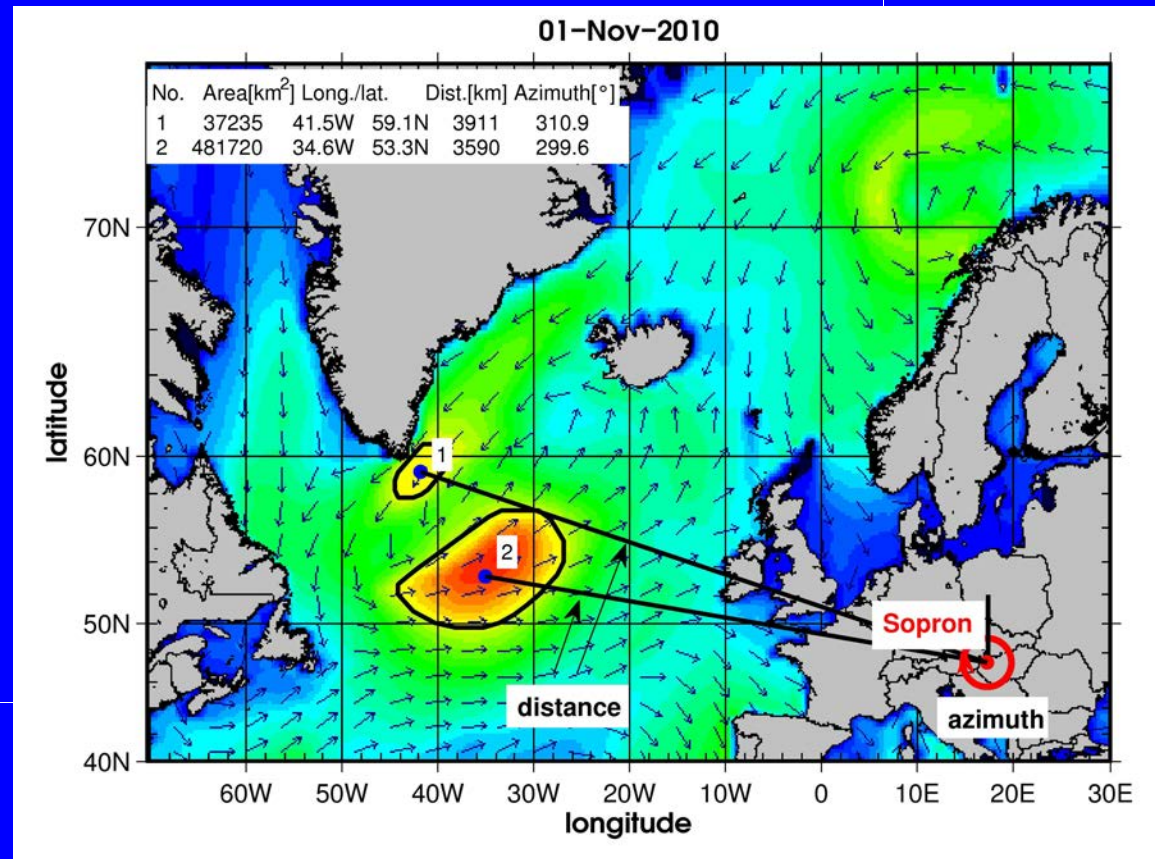


Identification of the sources of microseisms

- The origin of microseismic "noise" (Longuet-Higgins, 1950)
- Monitoring of ocean weather (e.g. Oceanweather Inc. USA)
- Parameterization of triggering events (storm zones)

- distance ($C(\varphi, \lambda)$)
- azimuth
- S area of storm zone ($H_{sw} > 6.5$ m)
- significant wave volume:

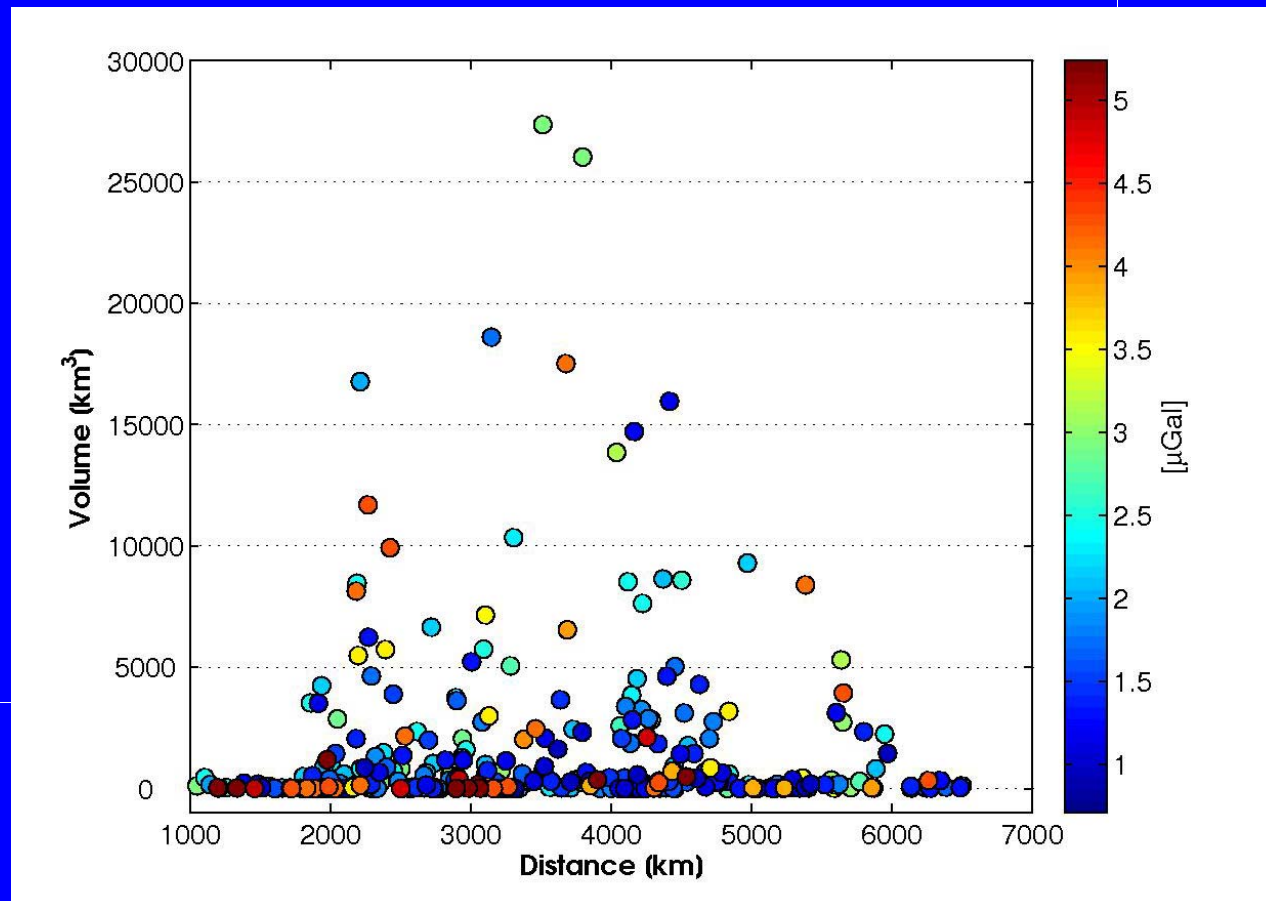
$$V_{sw} = \sum_{i=1}^M S_i(H_{sw})_i$$



Identification of the sources of microseisms

- Relations between the parameters of triggering events and the observed noise level

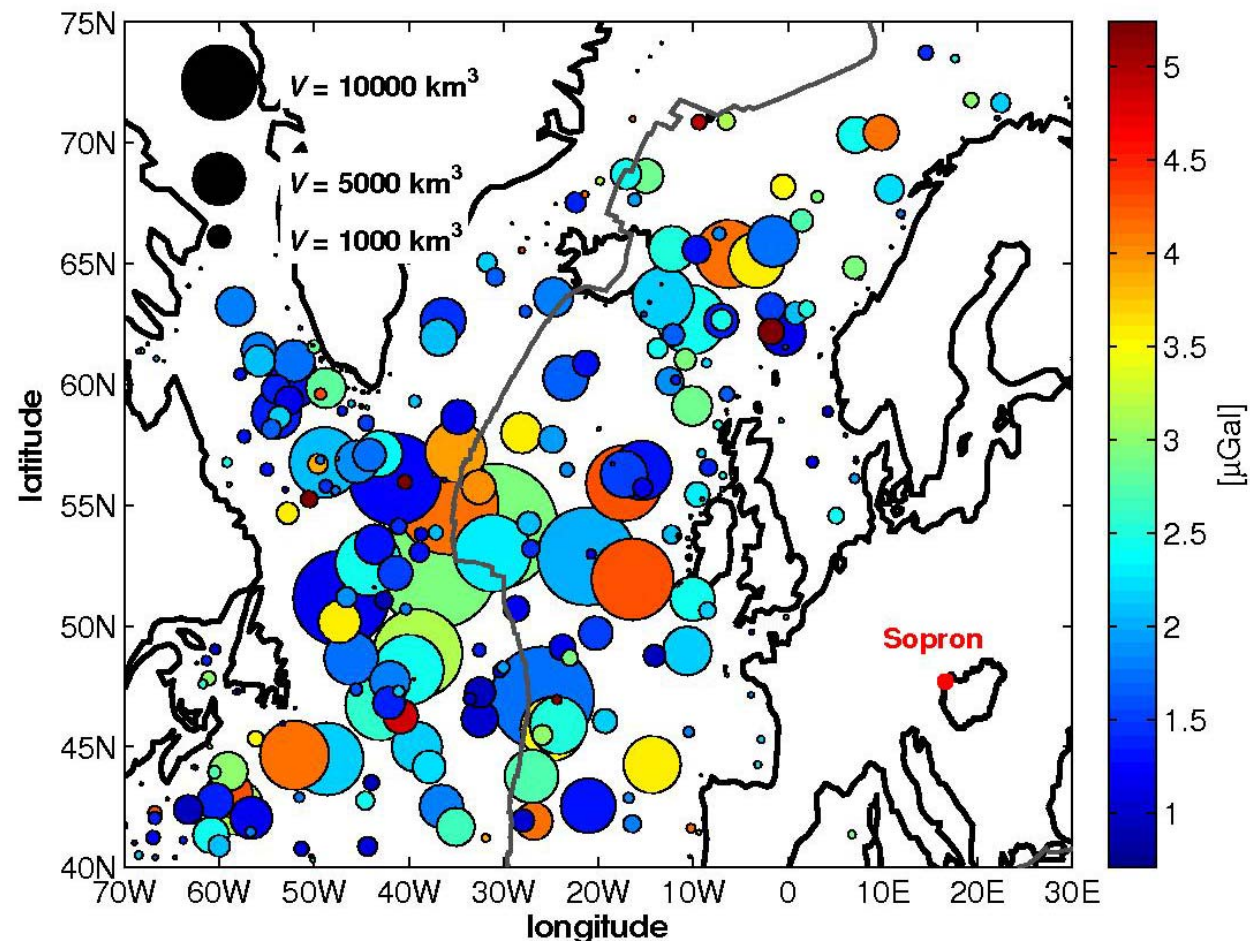
distance + V_{SW} vs. noise level



Identification of the sources of microseisms

- Relations between the parameters of triggering events and the observed noise level

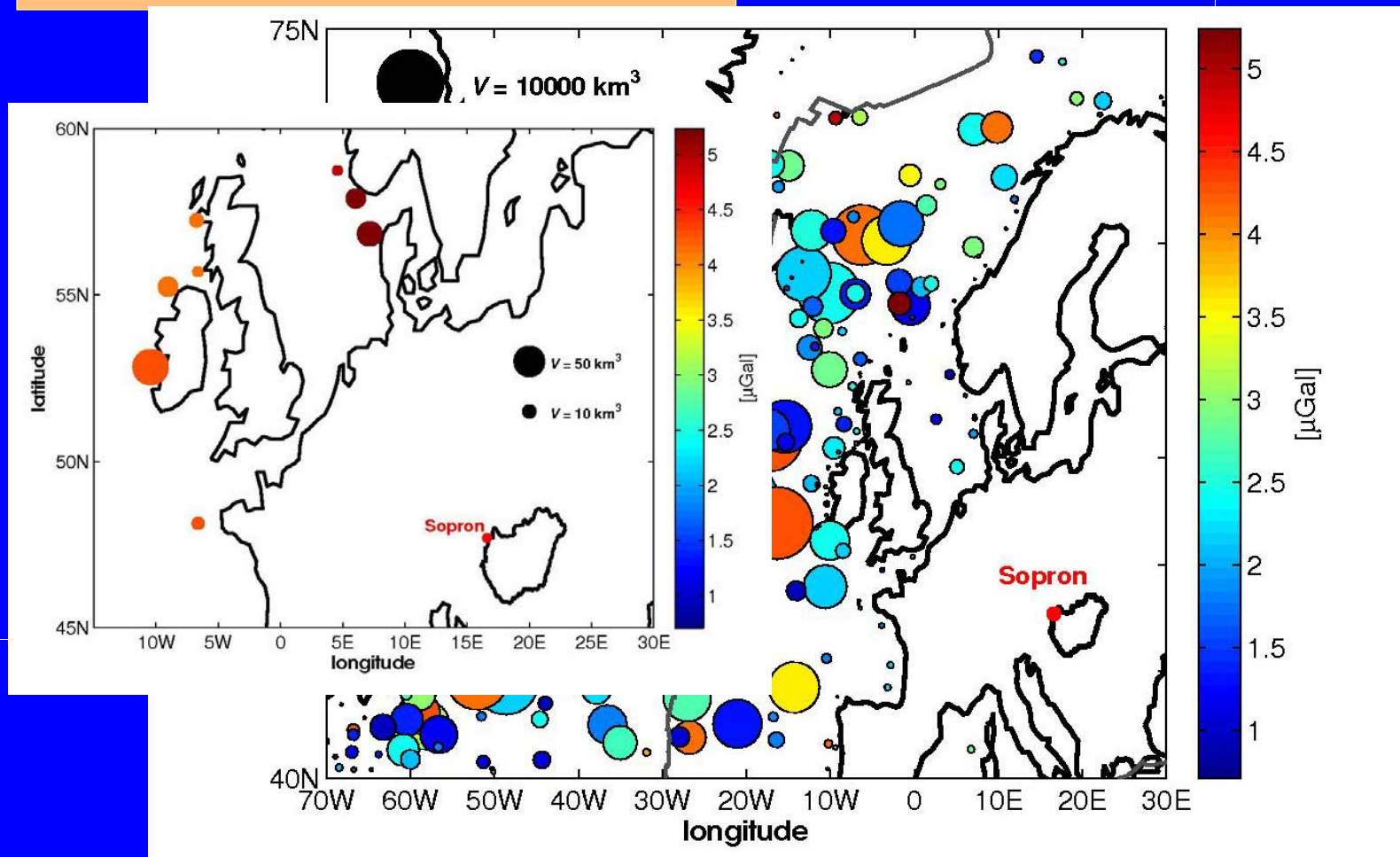
$C(\varphi, \lambda) + V_{SW}$ vs. noise level



Identification of the sources of microseisms

- Relations between the parameters of triggering events and the observed noise level

$C(\varphi, \lambda) + V_{SW}$ vs. noise level



Spectral analysis of 1 Hz records

- Co-located observations with spring and superconducting gravity meters in the Conrad Observatory (ZAMG), Austria
Time period: January 12. 2012 - May 2. 2013
Instruments: GWR SG025 (ZAMG)
Scintrex CG (Univ. Vienna)
LCR G220 (ELGI)
LCR G949 (MTA CSFK GGI)

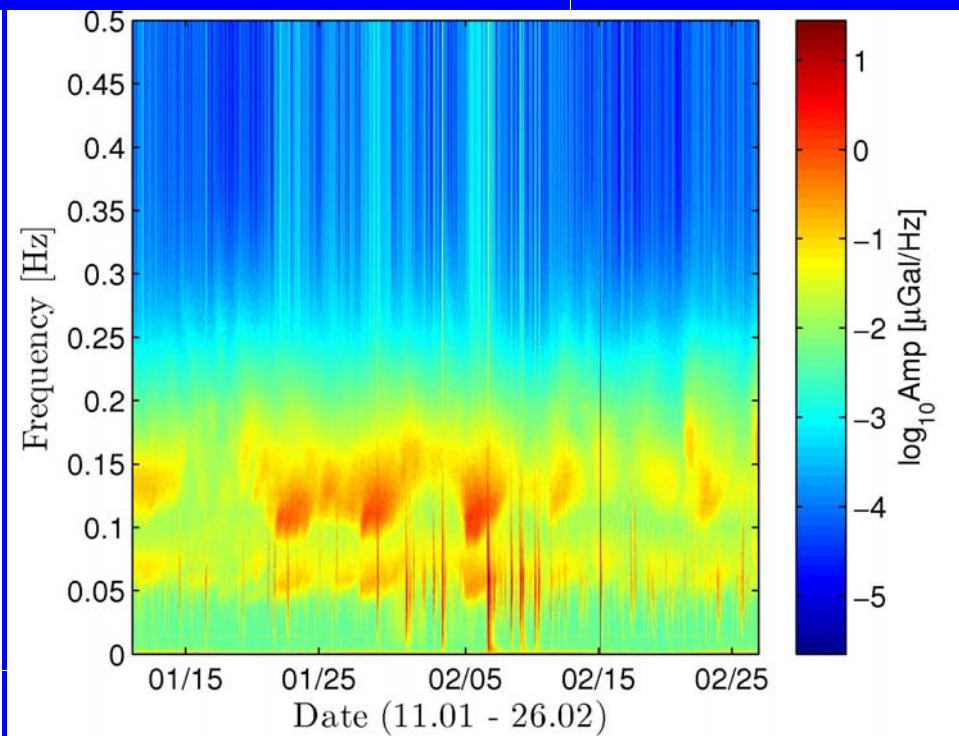
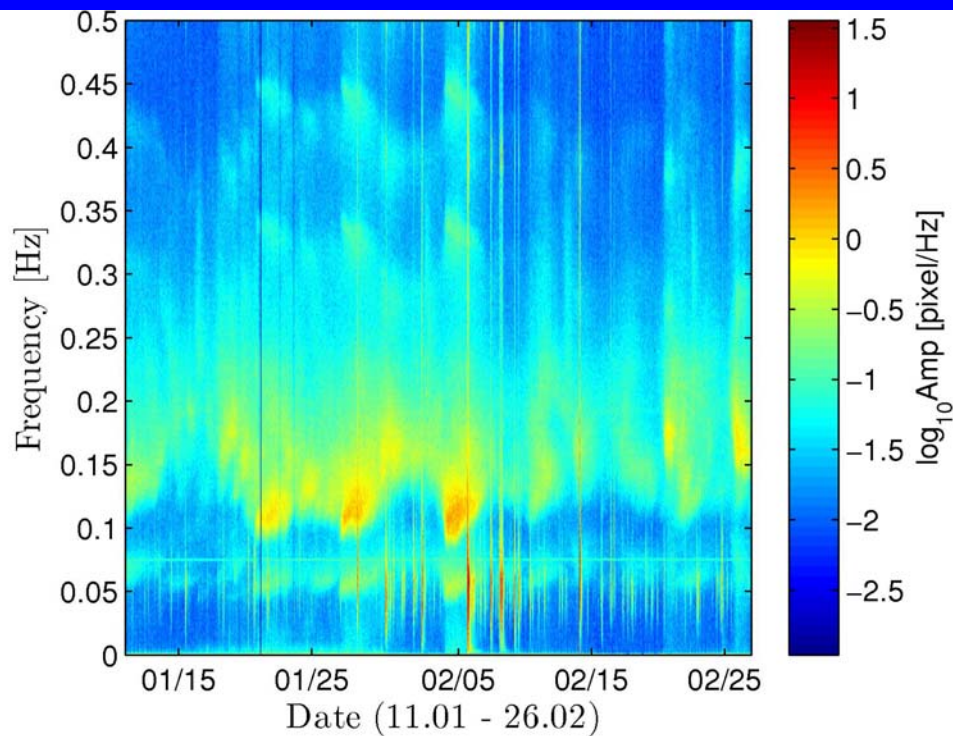
Spectral analysis of 1 Hz records

- Direct comparison of spectra of recorded time series

Dynamic spectra from 1 h long record segments

LCR G949

GWR SG025



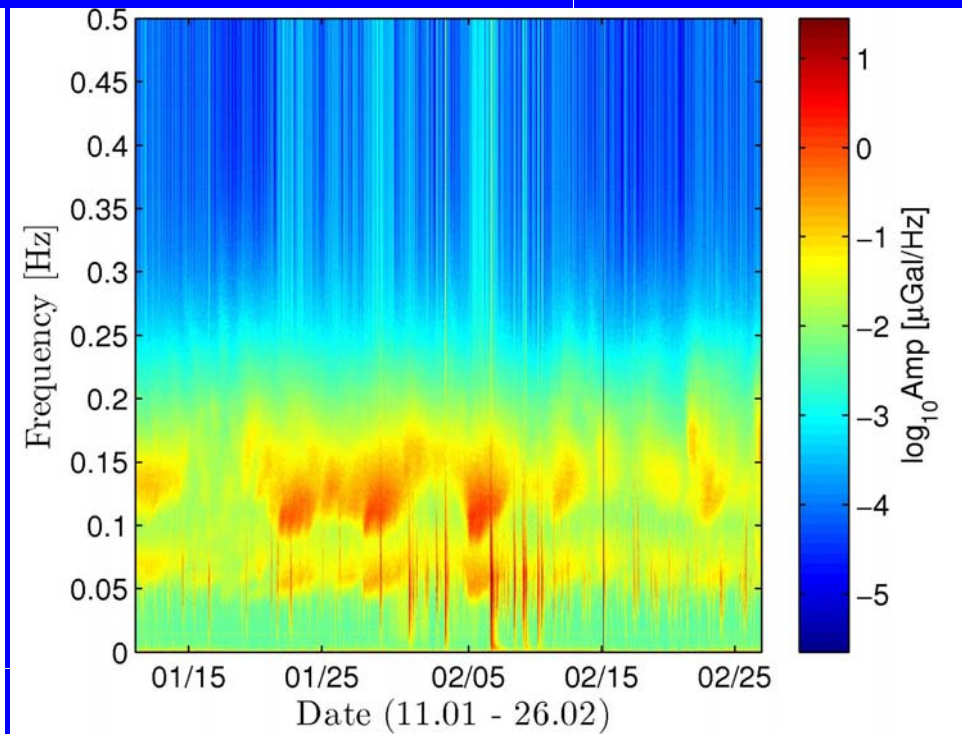
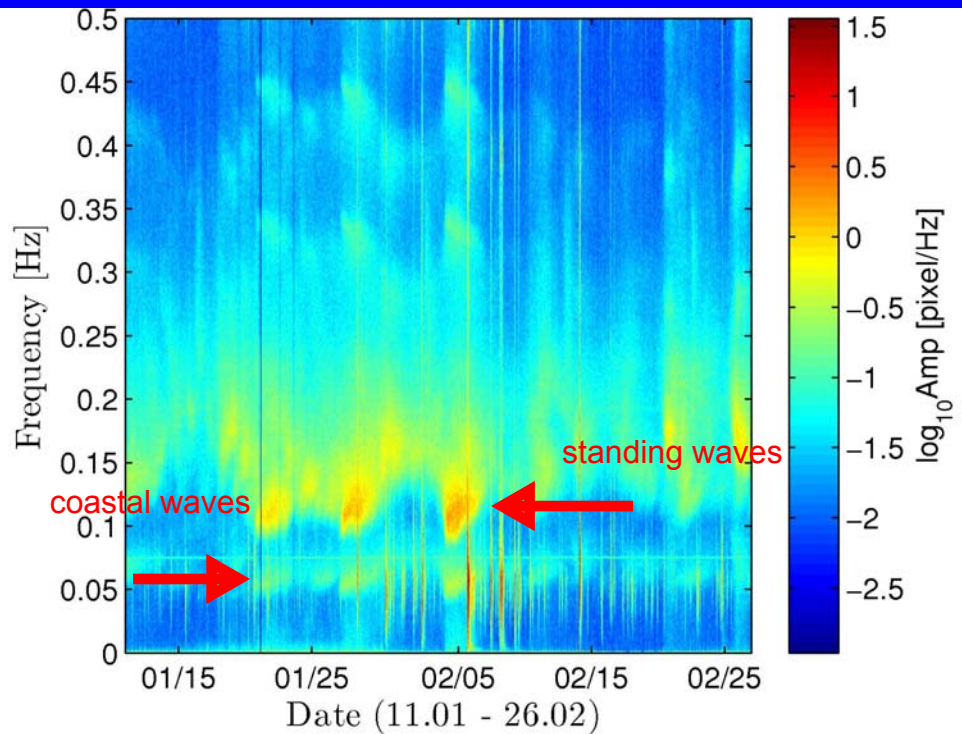
Spectral analysis of 1 Hz records

- Direct comparison of spectra of recorded time series

Dynamic spectra from 1 h long record segments

LCR G949

GWR SG025

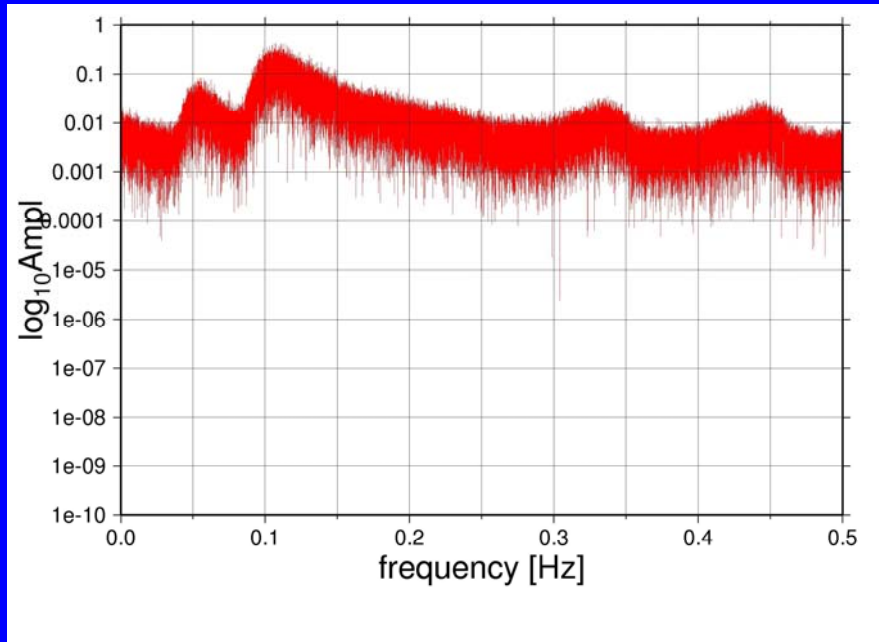


Spectral analysis of 1 Hz records

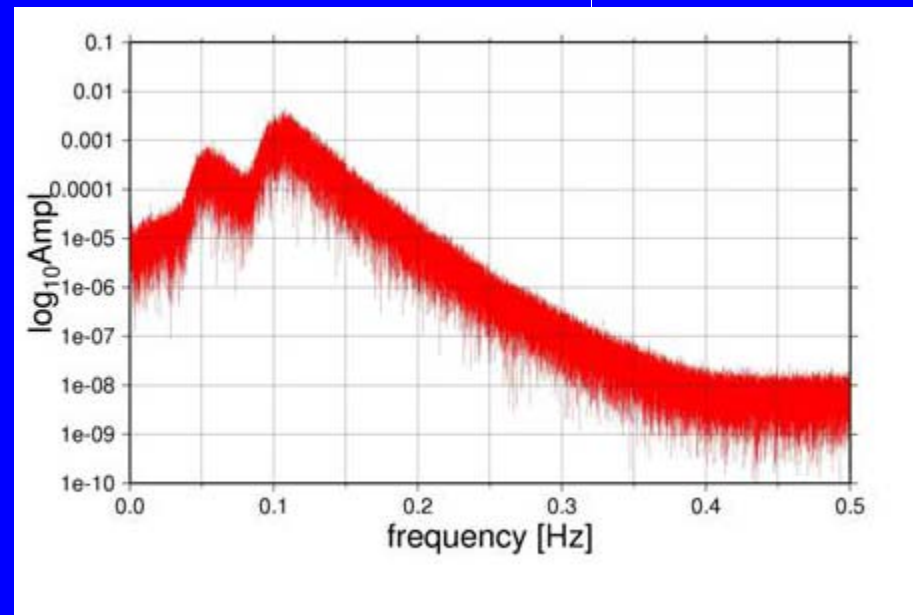
- Direct comparison of spectra of recorded time series

Spectra for the period 24.02.2013 - 26.02.2013 Conrad Observatory (Austria)

LCR G949



SG 025



Acknowledgements

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