

# TIGER's tail: Testing the no-hair theorem with black hole ringdowns.

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Meidam et al., Phys. Rev D 90 (2014) 064009

- *No-hair theorem*: Final stationary black hole vacuum state is determined only by its mass  $M$  and spin  $J$ .
- Quasinormal modes are characterized by mode frequencies and damping times  $\omega_{lm}(M, J)$  and  $\tau_{lm}(M, J)$ .
- Deviations from predicted mass and spin dependence can be written as:

$$\omega_{lm} = \omega_{lm}^{\text{GR}}(M, J)(1 + \delta\hat{\omega}_{lm}), \quad (1)$$

$$\tau_{lm} = \tau_{lm}^{\text{GR}}(M, J)(1 + \delta\hat{\tau}_{lm}) \quad (2)$$

- Use multiple detections from ET to test the no-hair theorem

## Previous work

- Two methods:

- ①  $M = 500 - 1000 M_{\odot}$ ,  $D_L = 6 \text{ Gpc}$  ( $z \sim 1$ ):

Determined measurability of  $\delta\hat{\omega}_{22}$ ,  $\delta\hat{\omega}_{33}$  and  $\delta\hat{\tau}_{22}$ .

$\Rightarrow$  Accuracies of a few percent for the first two, and about 10% for the third.

- ②  $M = 500 M_{\odot}$ , compare evidences for two models:

- model1:  $\{\delta\hat{\omega}_{22}, \delta\hat{\omega}_{33}, \delta\hat{\tau}_{22}\}$  free parameters,

- model2: parameters are zero, i.e. the GR prediction.

$\Rightarrow$  Deviation of a few percent in  $\delta\hat{\omega}_{22}$  could be discriminated from GR.

Gossan et al. Phys. Rev. D 85, 124056 (2014)

## Using multiple detections

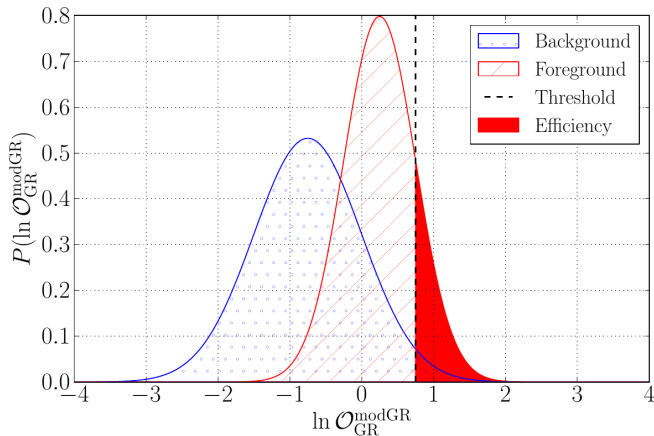
- Coalescence rates in the range  $500 - 1000 M_{\odot}$  estimated between one and a few thousand per year.
- Maximally exploit available set of detections by combining information from many sources.
  - ① Combining posteriors for  $\delta\hat{\omega}_{22}$ ,  $\delta\hat{\omega}_{33}$  and/or  $\delta\hat{\tau}_{22}$  not possible unless assuming GR is correct.
  - ② Bayesian model selection well suited.  
However, if  $\{\delta\hat{\omega}_{22}, \delta\hat{\omega}_{33}, \delta\hat{\tau}_{22}\}$  allowed to vary at the same time, **one may be penalized if correct model involves a smaller number of additional parameters.**

## Test Infrastructure for GEneral Relativity

- TIGER is designed and well tested for binary neutron star sources and advanced detectors.
- TIGER is well suited for low SNR and information from a population of sources can be combined.
- Simply put: TIGER compares evidences between two hypotheses:
  - $\mathcal{H}_{GR}$ : GR is correct
  - $\mathcal{H}_{\text{modGR}}$ : One or more parameters deviate from zero.
- Use odds ratio to compare models

$$\mathcal{O}_{GR}^{\text{modGR}} \equiv \frac{P(\mathcal{H}_{\text{modGR}}|d, I)}{P(\mathcal{H}_{GR}|d, I)} \quad (3)$$

## What are the odds?



## Hypotheses

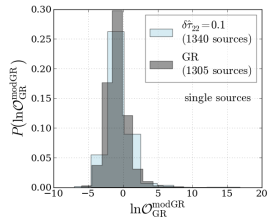
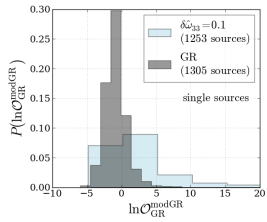
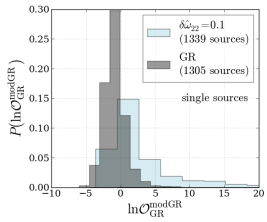
- The dimensionless parameterized deformations are  $\{\delta\hat{\omega}_{22}, \delta\hat{\omega}_{33}, \delta\hat{\tau}_{22}\}$ . Recall that

$$\omega_{lm} = \omega_{lm}^{\text{GR}}(M, J)(1 + \delta\hat{\omega}_{lm}),$$

$$\tau_{lm} = \tau_{lm}^{\text{GR}}(M, J)(1 + \delta\hat{\tau}_{lm})$$

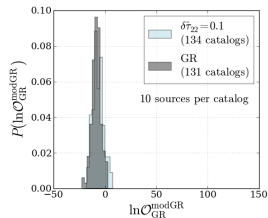
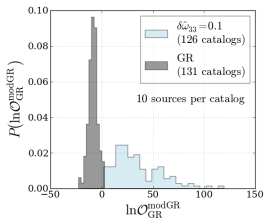
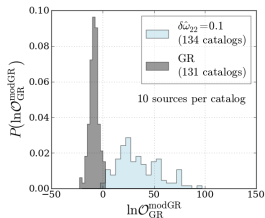
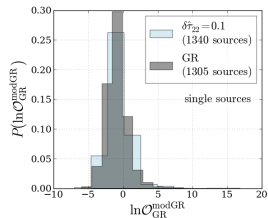
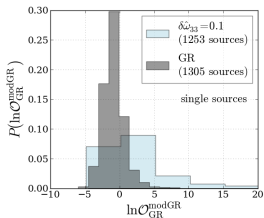
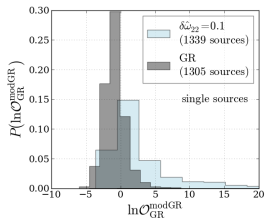
- The four cases analyzed in this work are:
  - ①  $(\delta\hat{\omega}_{22}, \delta\hat{\omega}_{33}, \delta\hat{\tau}_{22}) = (0.1, 0, 0)$ .
  - ②  $(\delta\hat{\omega}_{22}, \delta\hat{\omega}_{33}, \delta\hat{\tau}_{22}) = (0, 0.1, 0)$ .
  - ③  $(\delta\hat{\omega}_{22}, \delta\hat{\omega}_{33}, \delta\hat{\tau}_{22}) = (0, 0, 0.1)$ .
  - ④  $(\delta\hat{\omega}_{22}, \delta\hat{\omega}_{33}, \delta\hat{\tau}_{22}) = (0, 0, 0.25)$ .

# Single sources vs. combined sources

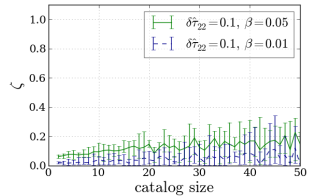
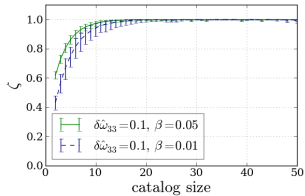
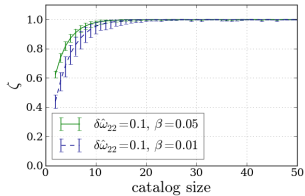




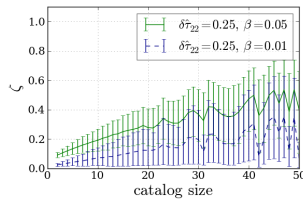
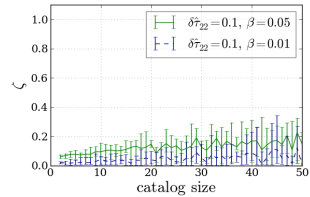
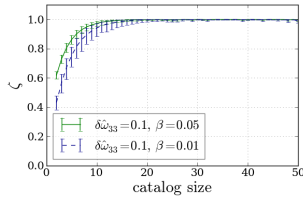
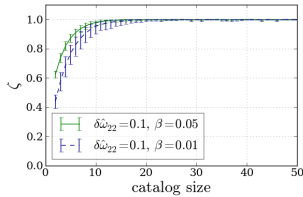
## Single sources vs. combined sources



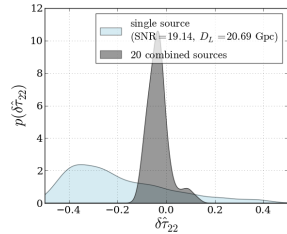
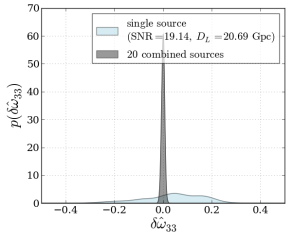
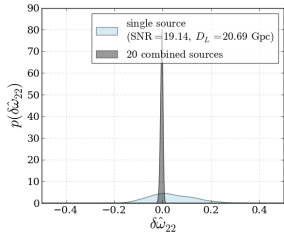
# Efficiency



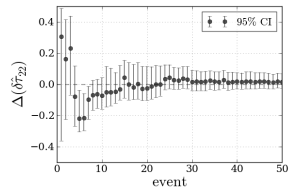
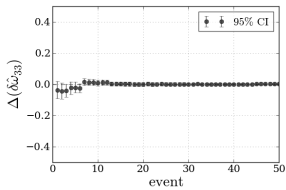
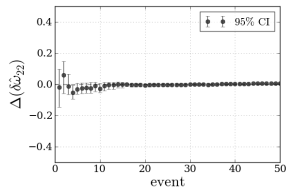
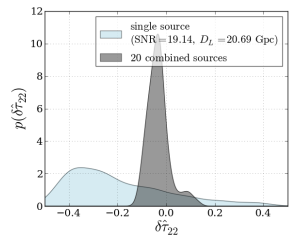
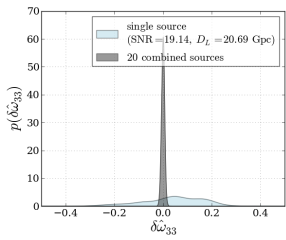
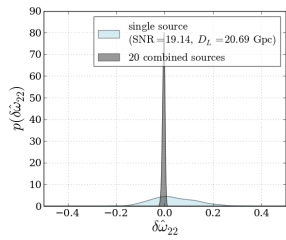
# Efficiency



# Parameter estimation



# Parameter estimation



## Summary

- Previous work has shown how deviation of a few percent in  $\omega_{22}$ ,  $\omega_{33}$  and  $\tau_{22}$  could be observed up to  $z \sim 1$  ( $D_L \sim 6$  Gpc).
- We used the TIGER framework to maximally exploit the available set of detections from ET in a low SNR scenario.
- We show that a deviation of a few percent can be observed up to  $z \sim 5$  ( $D_L \sim 50$  Gpc), when combining  $\mathcal{O}(10)$  sources.
- When GR is not in doubt, upper limits can be placed on deviations.



Thank you





- For violations of the no-hair theorem we found the following efficiencies:

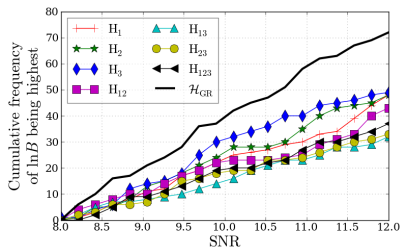
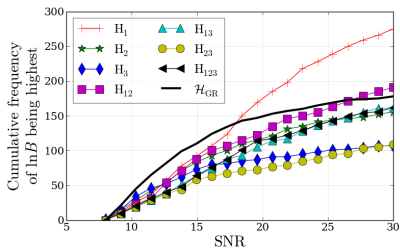
	single source	10 sources	50 sources
$\delta\hat{\omega}_{22} = 0.10$	$\sim 60 - 65\%$	$> 98\%$	$> 99\%$
$\delta\hat{\omega}_{33} = 0.10$	$\sim 60 - 65\%$	$> 95\%$	$> 99\%$
$\delta\hat{\tau}_{22} = 0.10$	$\sim 7 - 10\%$	$\sim 7 - 18\%$	$\sim 7 - 18\%$
$\delta\hat{\tau}_{22} = 0.25$	$\sim 10 - 14\%$	$\sim 12 - 30\%$	$\sim 20 - 70\%$

- When GR is not in doubt, the following constraints could be placed on the free parameters:

	single source	20 sources	50 sources
$\delta\hat{\omega}_{22}$	0.10	0.0051	$\sim 0.001$
$\delta\hat{\omega}_{33}$	0.13	0.0051	$\sim 0.001$
$\delta\hat{\tau}_{22}$	0.21	0.0480	$\sim 0.1$

## More advantages to using TIGER

- Combining sources greatly improves ability to test the no-hair theorem.
- Using multiple subhypotheses has a significant impact in finding deviations.



## More advantages to using TIGER

- As sources are combined, two observations can be made:
  - $H_{123}$  ( $\{\delta\hat{\omega}_{22}, \delta\hat{\omega}_{33}, \delta\hat{\tau}_{22}\}$  are left free) is deprecated compared to some others!
  - The correct hypothesis is not always the most dominant one!

