

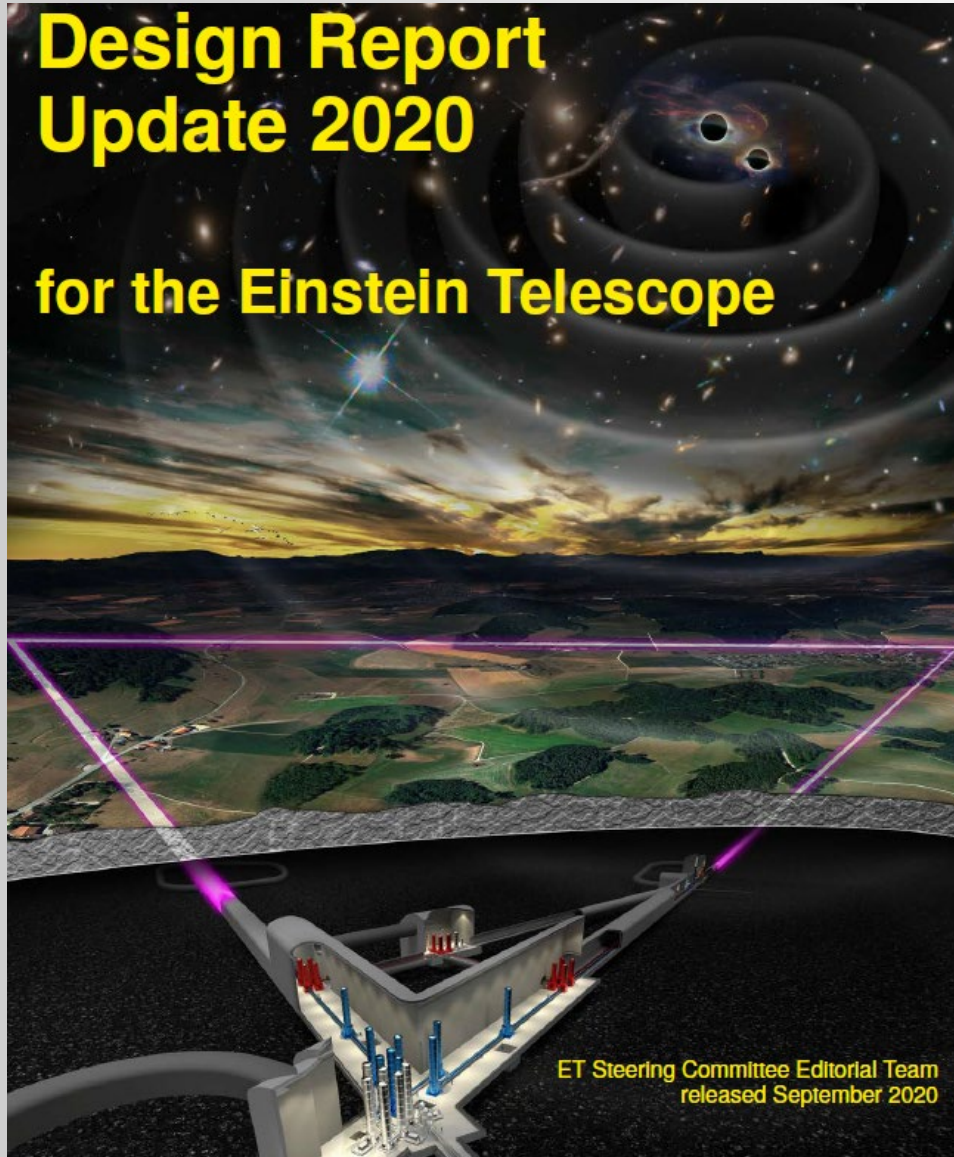
# Stabilized High-Power Lasers for the Einstein Telescope

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16 June 2021 (ET Laser Work Package kick-off meeting)



# most current ET design information



- most current ET design information can be found in

**Design Report Update 2020 for the Einstein Telescope**

- info concerning the stabilized laser system in

**Chapter 6.4: Light sources**

url:<https://apps.et-gw.eu/tds/?content=3&r=17245>

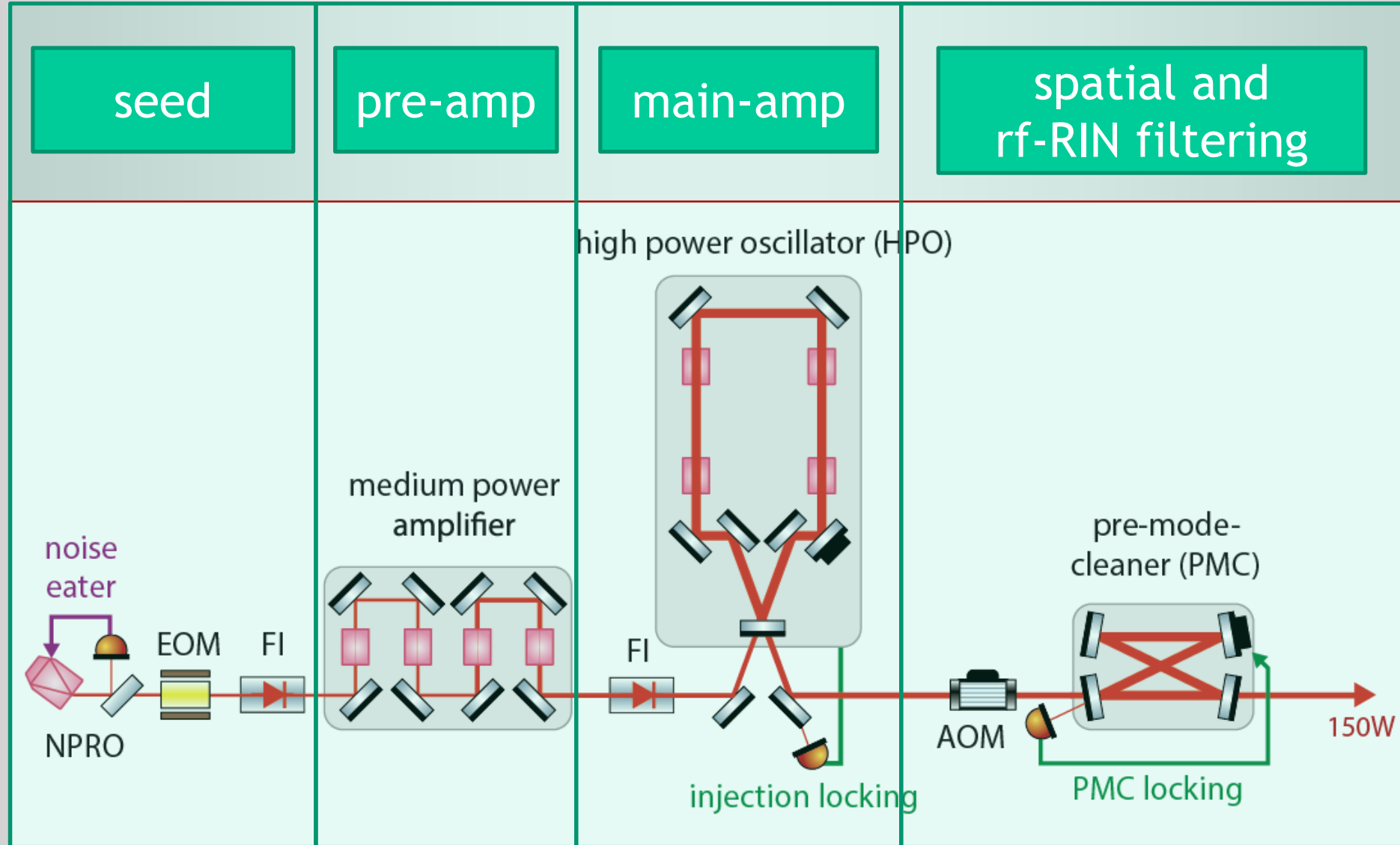


# General Requirements

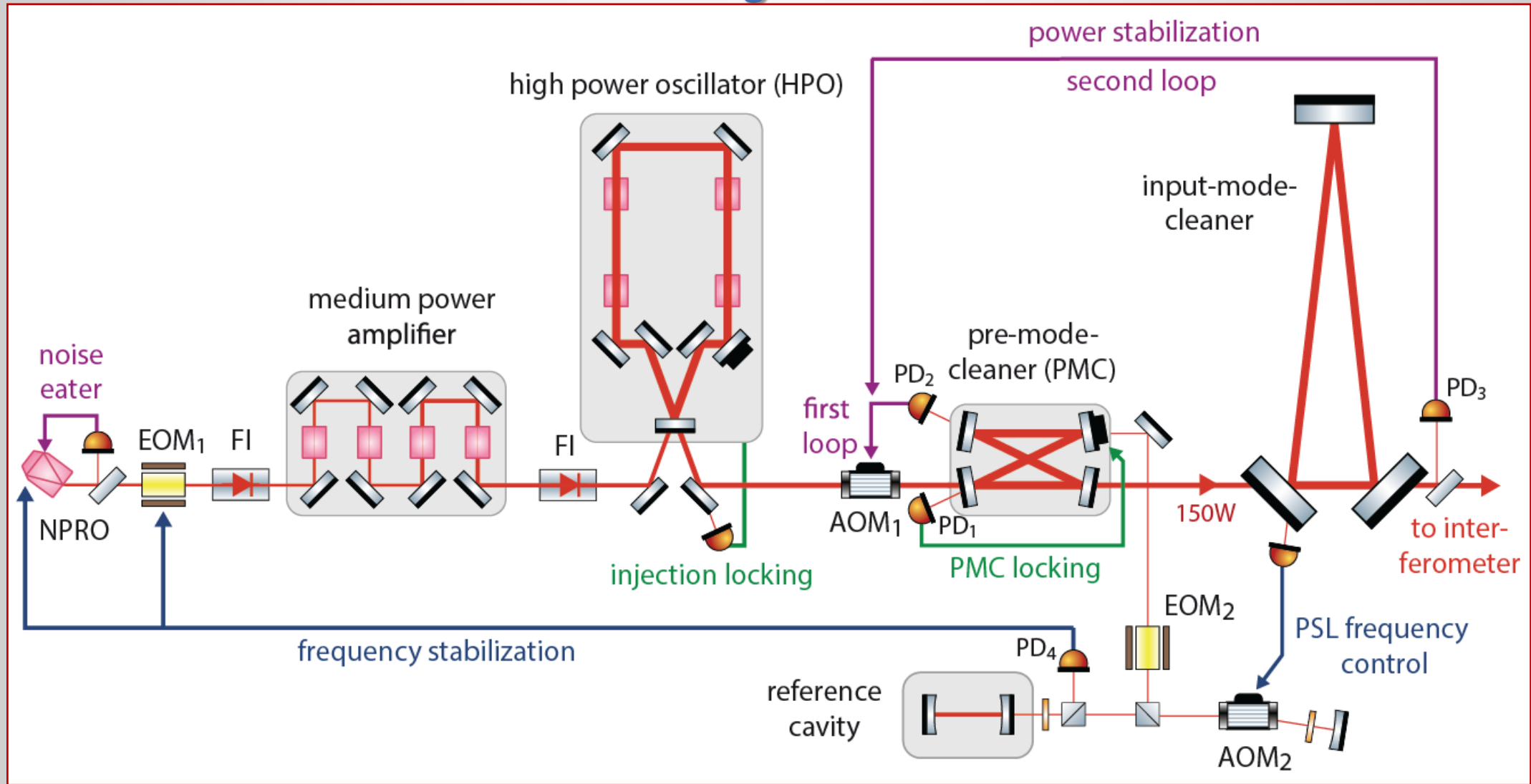
- **700 Watt power at 1064nm** (cw, single-frequency, linearly-polarized)
  - To enable high circulating power in interferometer arms
- **5 Watt power at 1550nm** (cw, single-frequency, linearly-polarized)
- **High spatial purity and low beam jitter**
  - Good coupling to input-mode-cleaner
  - Low shot noise on sensors for laser and input-mode-cleaner stabilization
- **Low free running noise** (in laser power and frequency)
  - Acceptable stabilization effort (loop gain and cross-couplings)
- **Low-noise sensors** (for laser power and frequency )
  - To achieve required stability for light entering the input mode-cleaner
- **Fast actuators** (for laser power and frequency) with large range
  - To allow for required loop gain in stabilization control loops
- **High robustness and reliability** with low maintenance requirements



# Typical GWD Laser Design (example aLIGO)



# Advanced LIGO Pre-Stabilized High-Power Laser





# Advanced LIGO Pre-Stabilized High-Power Laser



# Laser Requirements for ET

- cw, single-frequency, linearly-polarized with 700W @ 1064 and 5W @ 1550nm
- less than 10% higher order spatial mode content
- polarization purity > 1/10
- low free-running noise (equal or better than aLIGO and Advanced Virgo lasers)
  - power noise
  - frequency noise
  - beam pointing
  - polarization fluctuations
- power and frequency actuator
  - range larger than free running noise
  - speed sufficient to achieve  $RIN < 2 \cdot 10^{-10} 1/\sqrt{\text{Hz}}$  and  $\Delta\nu < 10 \mu\text{Hz}/\sqrt{\text{Hz}}$
- low drift of DC parameters and stable noise level
- robust and reliable 24/7 operation for years
- digital monitoring and control capabilities



# Stabilization Requirements

- **Filter Resonator (pre-modecleaner)**
  - higher order spatial mode reduction
  - beam pointing reduction
  - rf-power noise reduction
- **power noises sensor**
  - with shot-noise limited performance equivalent to 10 Watt
- **power stabilization feed-back control system**
  - relative electronic input noise of less than  $10^{-10} 1/\sqrt{\text{Hz}}$
  - high enough unity-gain frequency to achieve  $RIN < 2 \cdot 10^{-10} 1/\sqrt{\text{Hz}}$  at 10 Hz
- **frequency stabilization feed-back control system**
  - high dynamic range of more than 9 orders of magnitude at 10 Hz
  - high enough unity-gain frequency to achieve  $\Delta\nu < 10 \mu\text{Hz}/\sqrt{\text{Hz}}$
- **digital control and monitoring system**





# Competences Required in the ET Laser Group

- design and fabrication of
  - **lasers** at 1064 nm, 1550 nm and  $2\mu\text{m} - 2.1\mu\text{m}$
  - vacuum compatible **electro-optical modulators and Faraday Isolators** with high optical quality for high power levels
  - **power noise sensors** with noise below the shot-noise of several Watts
  - fast vacuum compatible **power actuator** for high power levels
- simulation and design of feed-back control systems
  - low noise **coherent beam combination** at 1064 nm
  - **noise analysis** of stabilized laser system including cross couplings
  - fast low-noise analog **feed-back control electronic**
- digital control and monitoring

