

Einstein Telescope

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On behalf of the ET steering committee

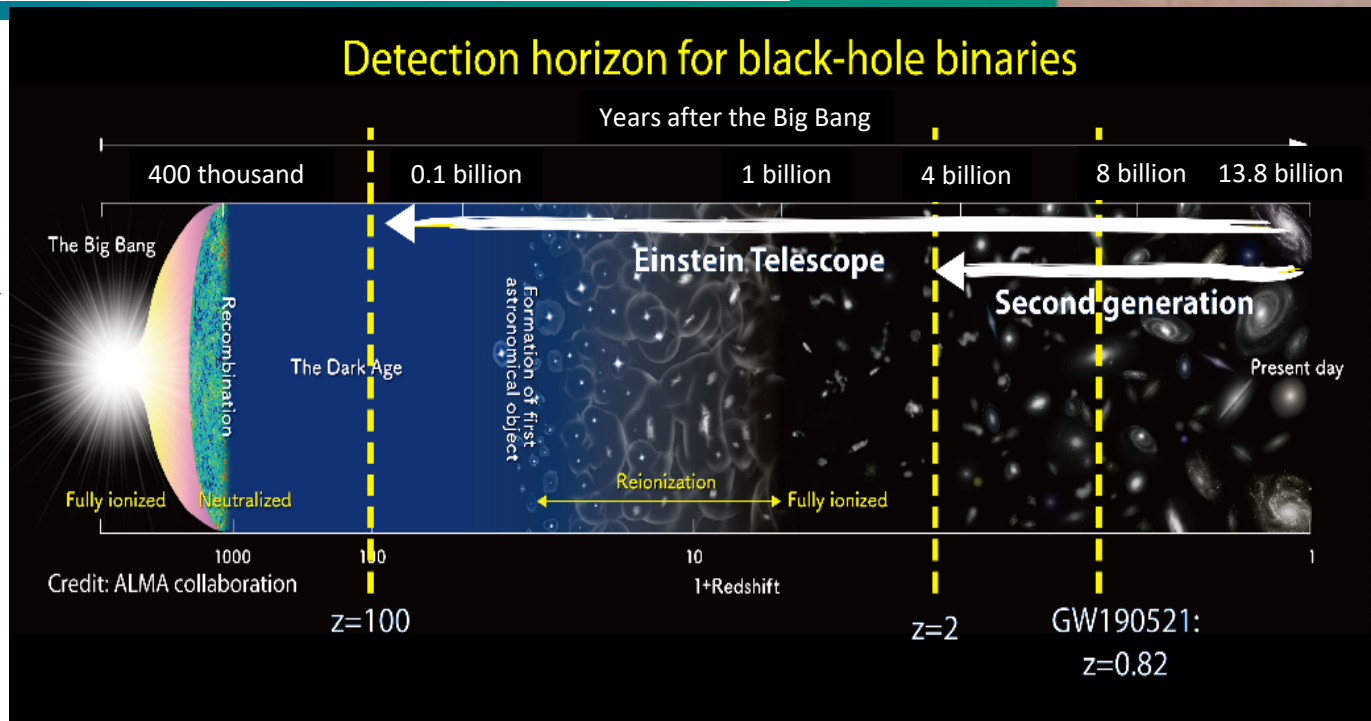
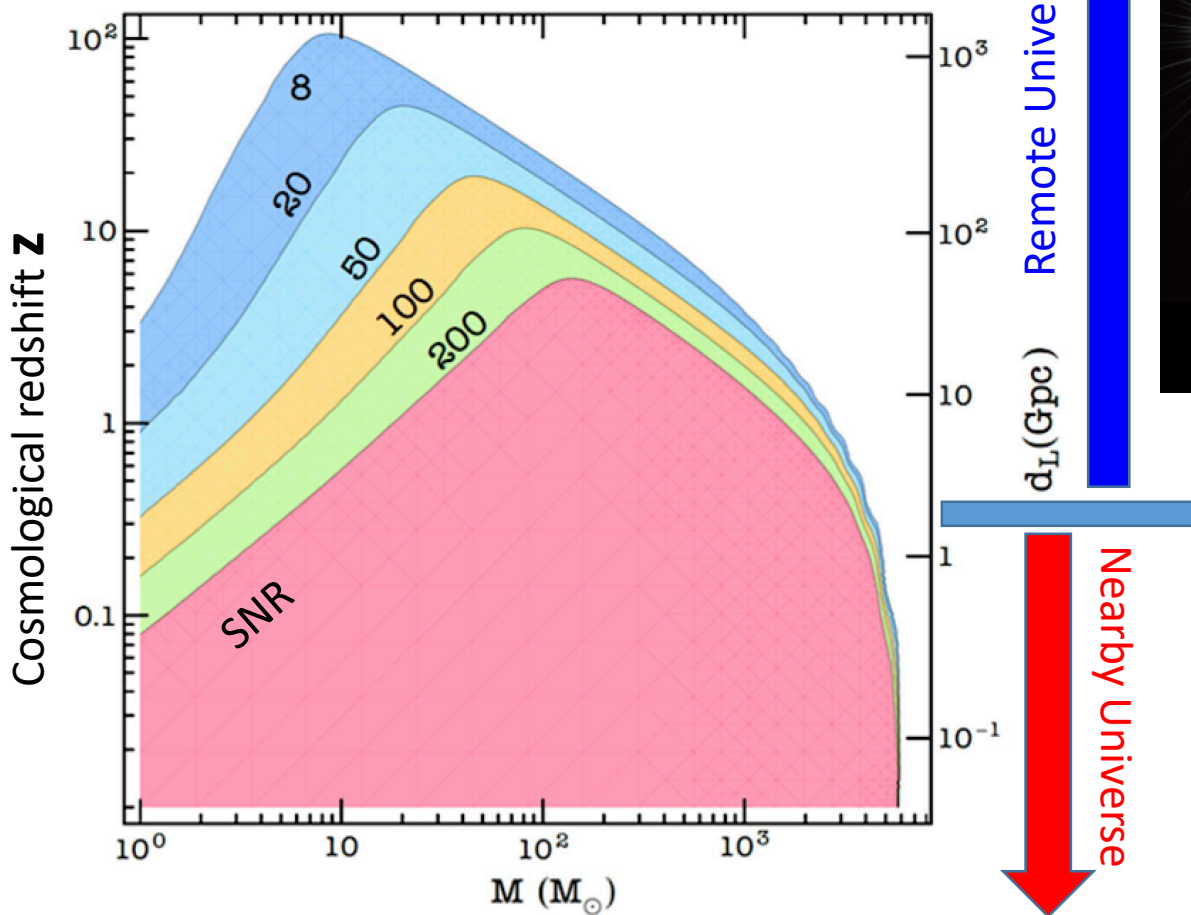
ET-0436A-21

What is Einstein Telescope (ET)?



- ET is the project aiming to realise the European 3rd Generation Gravitational Wave observatory
- ET has been a pioneer idea that defined the concept of 3rd generation GW observatory:
 - A sensitivity at least 10 times better than the (nominal) advanced detectors on a large fraction of the (detection) frequency band
 - Wideband (possibly wider than the current detectors) accessing the frequency band below 10Hz
 - High reliability and improved observation capability
- ET has a long and important history that formed first the ET community and now the ET project
- ET is now becoming also a (formal) scientific collaboration

- ET is the pioneer project of the 3G GW observatory in Europe
- Few words on ET science hereafter



The combination of

- distances and masses explored
- number of detections
- detections with very high SNR

will provide a wealth of data expected to generate **revolutions in astrophysics, cosmology and fundamental physics**

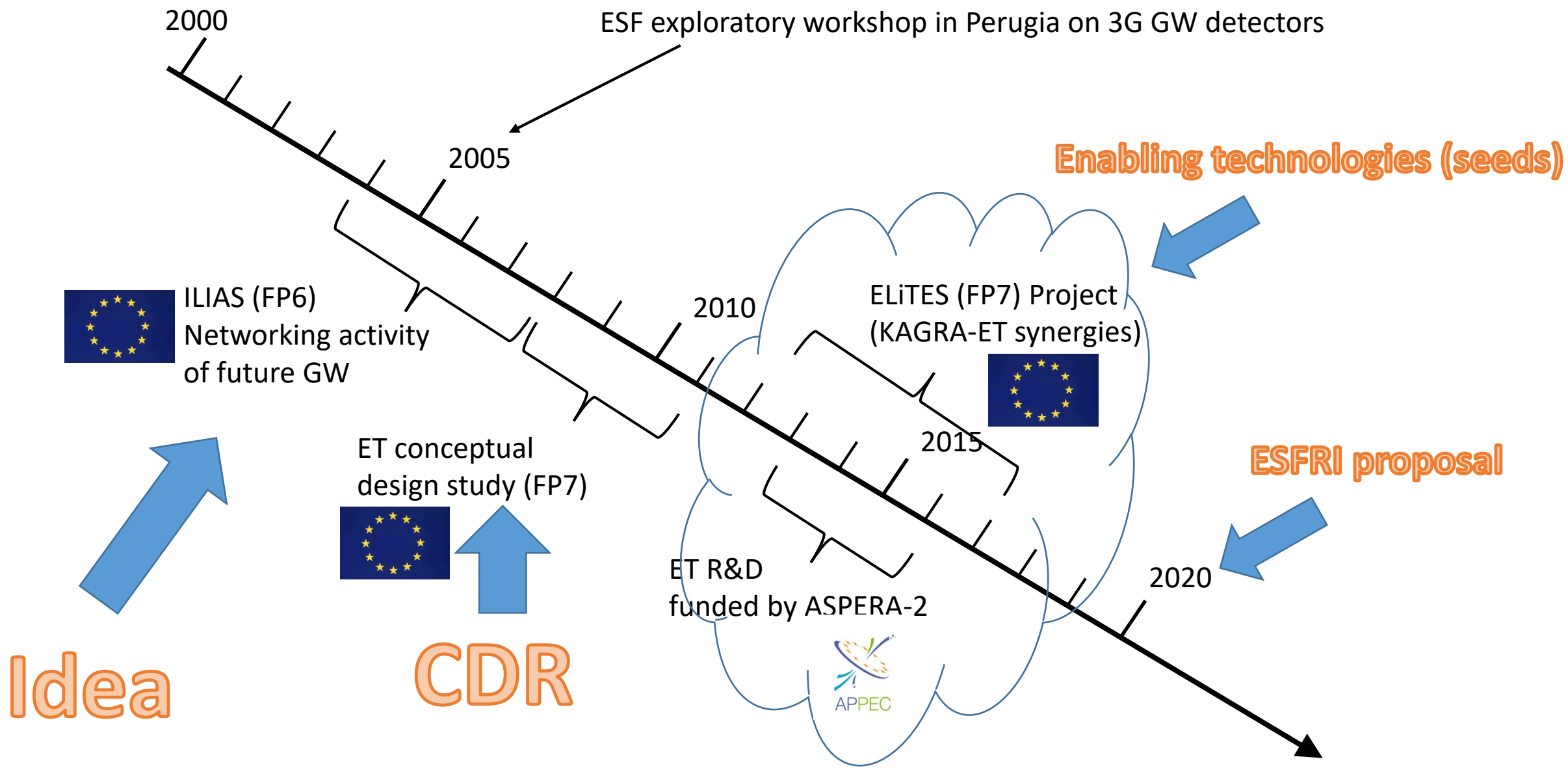
ASTROPHYSICS

- **Black hole properties**
 - origin (stellar vs. primordial)
 - evolution, demography
- **Neutron star properties**
 - interior structure (QCD at ultra-high densities, exotic states of matter)
 - demography
- **Multi-band and -messenger astronomy**
 - joint GW/EM observations (GRB, kilonova,...)
 - multiband GW detection (LISA)
 - neutrinos
- **Detection of new astrophysical sources**
 - core collapse supernovae
 - isolated neutron stars
 - stochastic background of astrophysical origin

FUNDAMENTAL PHYSICS AND COSMOLOGY

- **The nature of compact objects**
 - near-horizon physics
 - tests of no-hair theorem
 - exotic compact objects
- **Tests of General Relativity**
 - post-Newtonian expansion
 - strong field regime
- **Dark matter**
 - primordial BHs
 - axion clouds, dark matter accreting on compact objects
- **Dark energy and modifications of gravity on cosmological scales**
 - dark energy equation of state
 - modified GW propagation
- **Stochastic backgrounds of cosmological origin**
 - inflation, phase transitions, cosmic strings

ET long path



ESFRI Roadmap



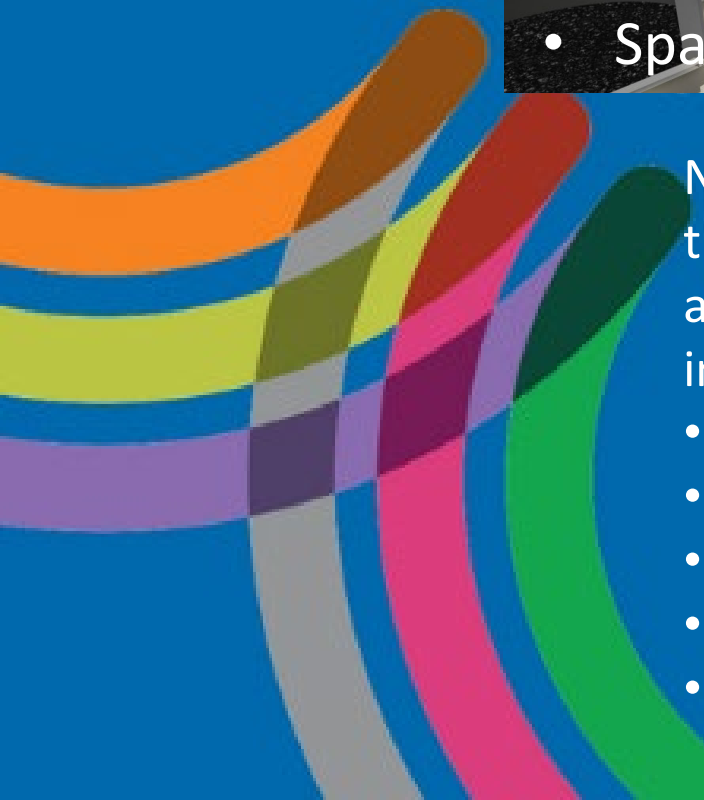
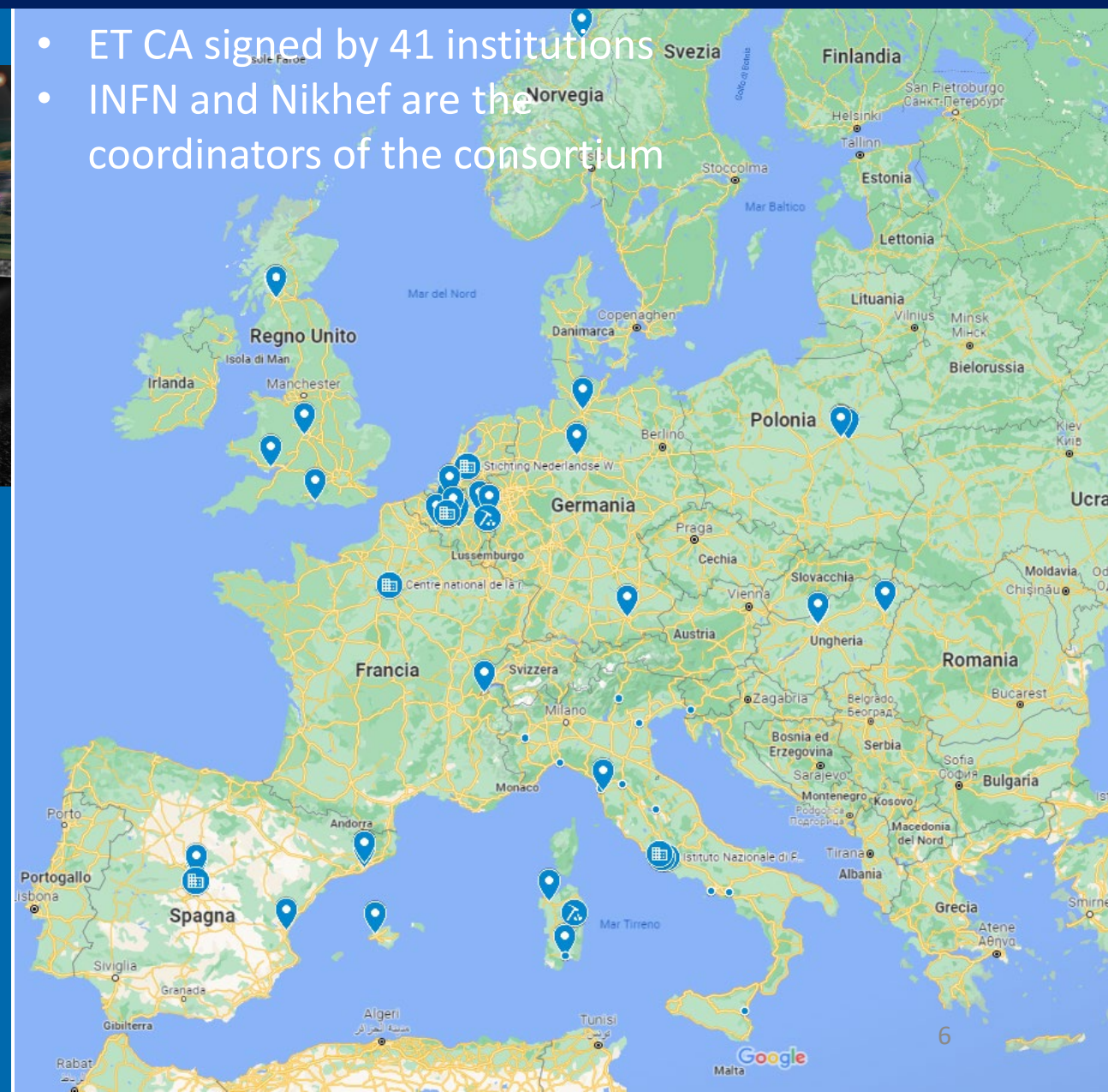
Proposal submitted by:

- Italy (Lead Country)
- Belgium
- Netherlands
- Poland
- Spain

Now in the project and in the collaboration activities also agencies or institutions belonging to:

- France
- Germany
- Hungary
- Switzerland
- UK

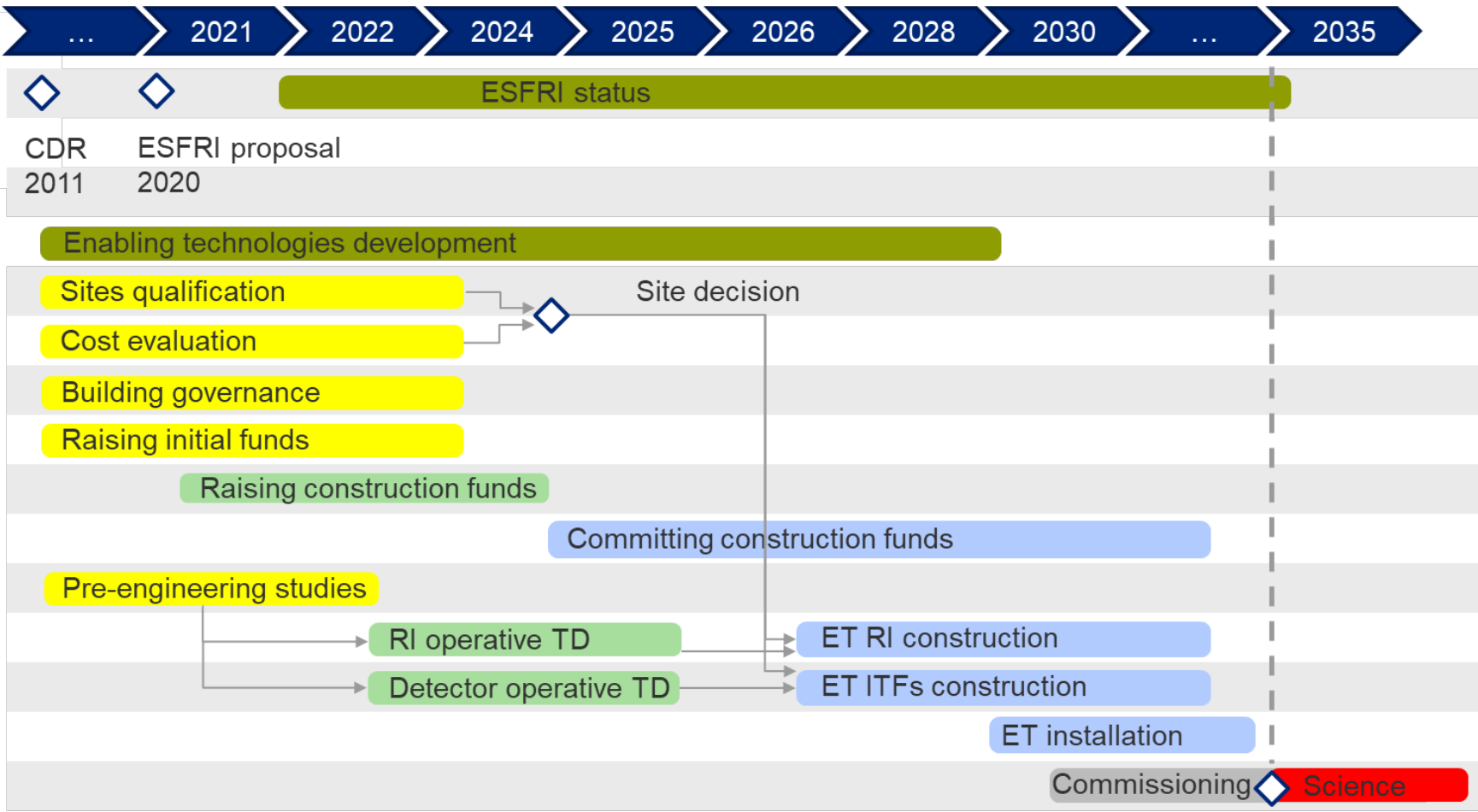
- ET CA signed by 41 institutions
- INFN and Nikhef are the coordinators of the consortium



ET timeline

- ET timeline presented to ESFRI
 - As expected, the ESFRI approval boosted the activities at all the levels:

* Tentative schedule

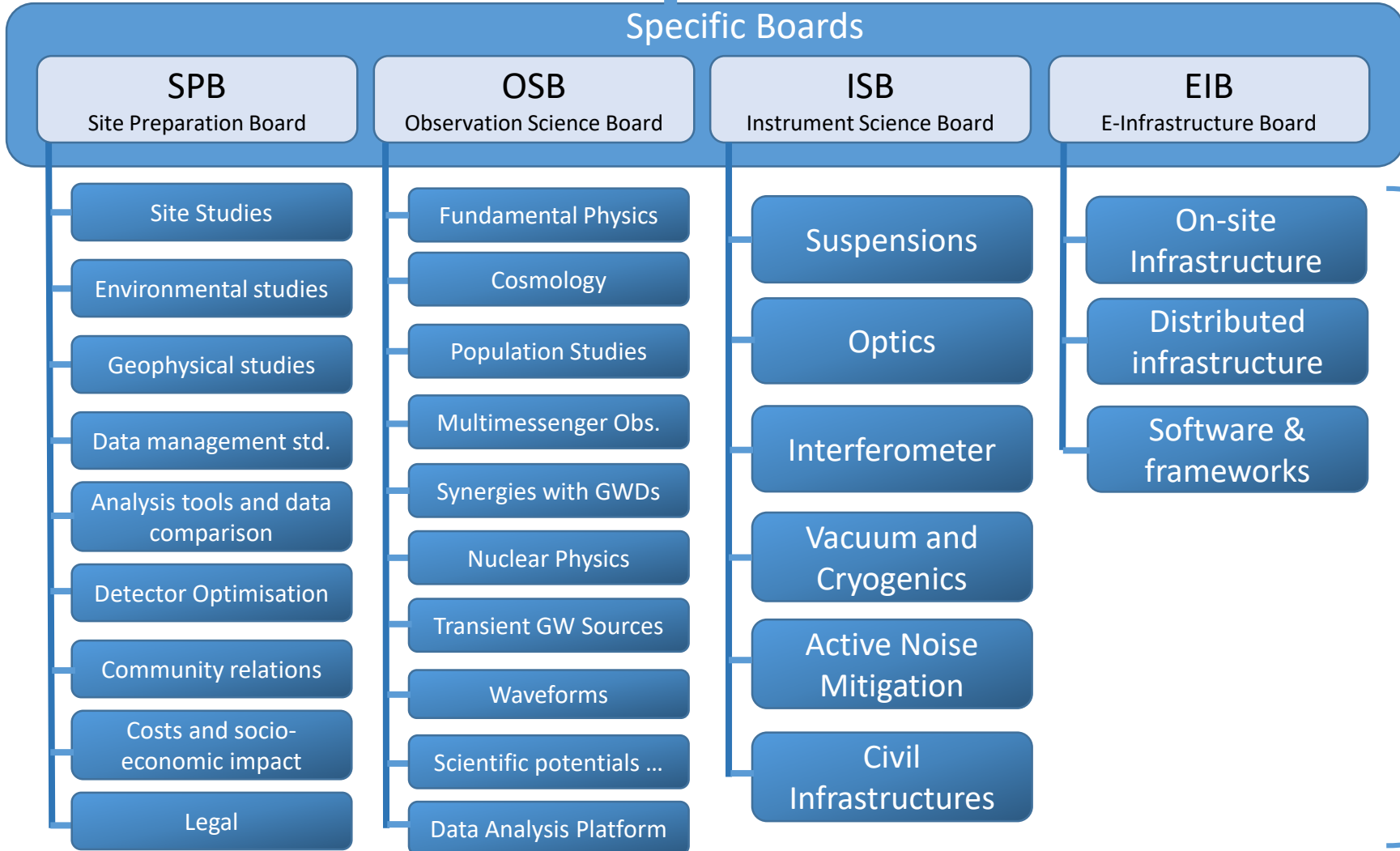


- Scientists
- Agencies
- Governments

ESFRI Phases: Design Preparatory Implementation Operation

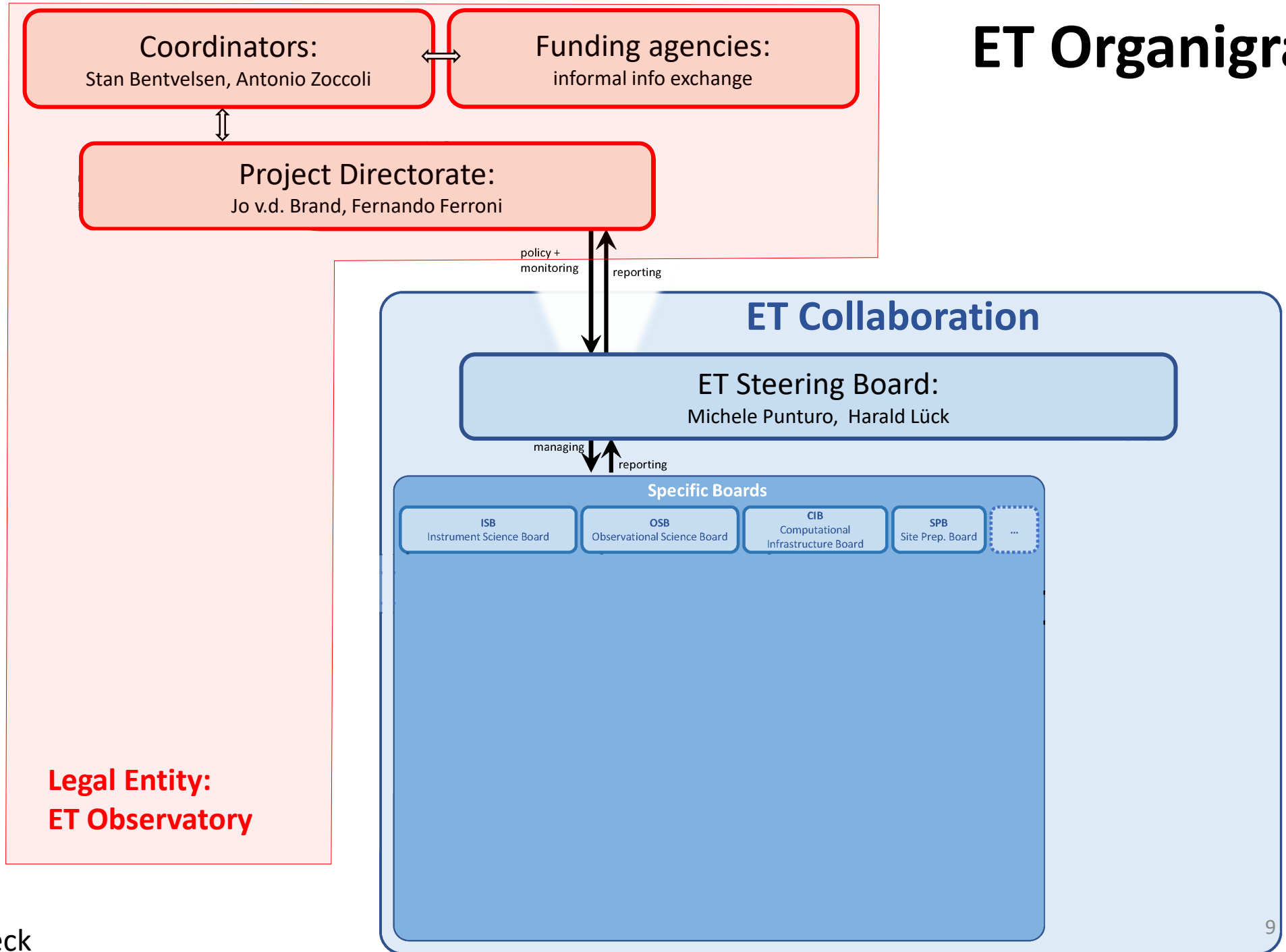


<http://www.et-gw.eu/index.php/et-steering-committee>



The proto-collaboration is currently the most organised component of the ET project

Divisions

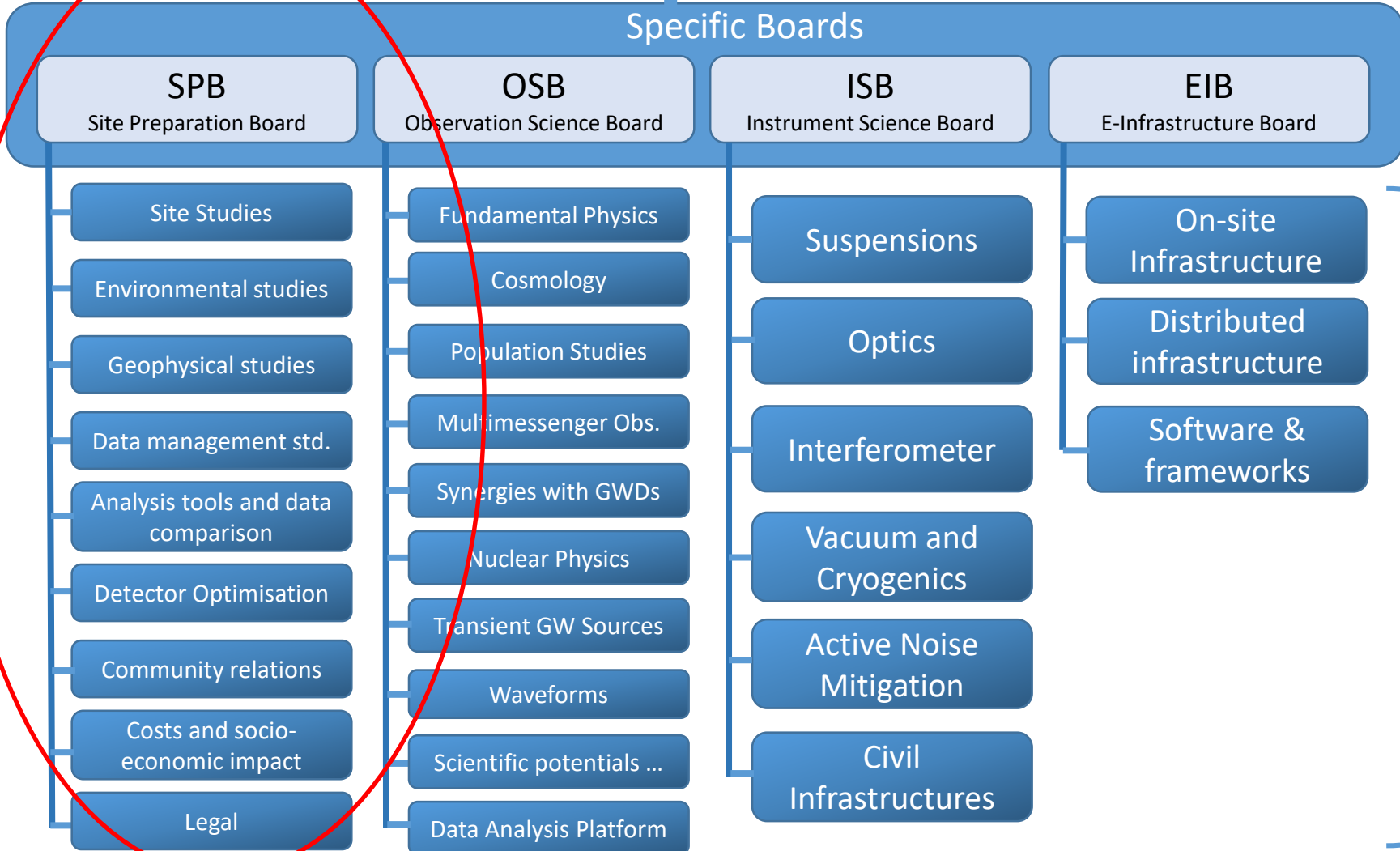


ET collaboration current organisation



ET Steering Committee

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Divisions

ET site(s)

- Currently there are two sites, in Europe, candidate to host ET:
 - The Sardinia site, close to the Sos Enattos mine
 - The EU Regio Rhine-Meuse site, close to the NL-B-D border
- A third option in Saxony (Germany) is under discussion, ~~but still too preliminary to be a candidate~~



SPB: ET sites under characterisation



Euregio Meuse-Rhine

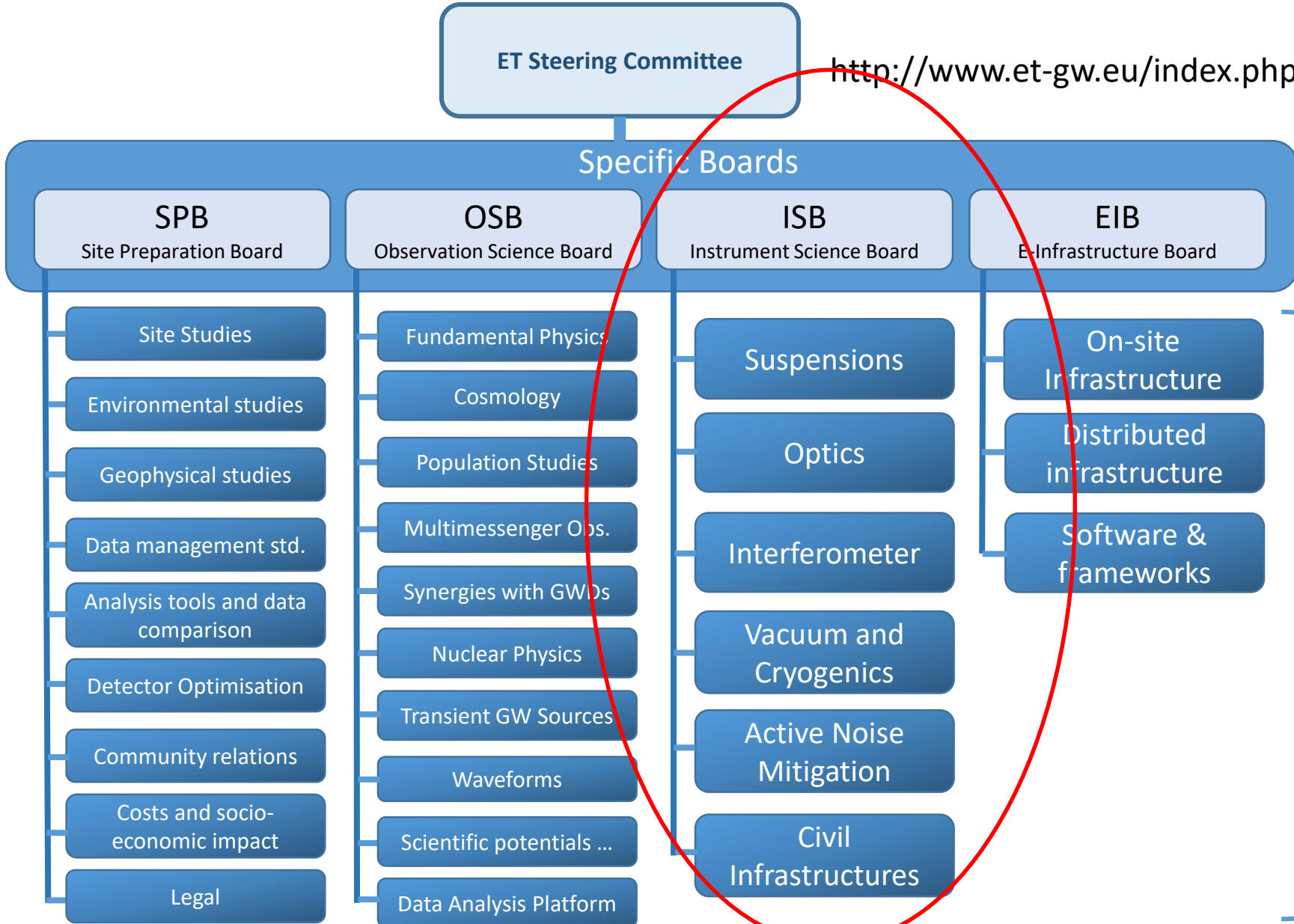
- A 250-m deep borehole has been excavated and equipped
 - Seismic data under acquisition and analysis
- 3-5 other boreholes expected
- Extensive active and passive site characterisation with sensor arrays in 2021
- Good seismic noise attenuation given by the particular geological structure
- Characterisation funded through Interreg grants

Sardinia

- Long standing characterisation of the mine in one of the corners continuing
 - Seismic, magnetic and acoustic noise characterisation ongoing at different depth in the mine
- Underground laboratory under construction (SarGrav)
- Two ~290m boreholes have been excavated, equipped and data taking is ongoing
- Intense & international surface investigations programme in Summer/Fall 2021
- Characterisation funded on regional and national funds

Link with the APOGEIA initiative (Stavros Katsanevas)?

ET collaboration current organisation

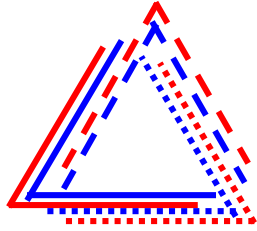


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Divisions

- The multi-interferometer approach asks for two parallel technology developments:



• **ET-LF:**

- Underground
- Cryogenics
- Silicon (Sapphire) test masses
- Large test masses
- New coatings
- New laser wavelength
- Seismic suspensions
- Frequency dependent squeezing

• **ET-HF:**

- High power laser
- Large test masses
- New coatings
- Thermal compensation
- Frequency dependent squeezing

Advanced detectors and their development programmes are a crucial de-risking factor for ET-HF

Evolved laser technology

Evolved technology in optics

Highly innovative adaptive optics

High quality opto-electronics and new controls

- Challenging engineering
- New technology in cryo-cooling
- New technology in optics
- New laser technology
- High precision mechanics and low noise controls
- High quality opto-electronics and new controls

Challenging engineering Large Underground Infrastructures

Technical challenges on:

Self-induced acoustic, seismic and e.m. noises

Temperature and Humidity stabilisation

Safety

New challenges in optics (and material science)

- ET-LF requires new materials for the large mirrors in order to operate at cryogenic temperature:
 - Silicon (Sapphire) mirrors:
 - Challenges on size and geometry, polishing, thermo-mechanical properties (thermal noise), optical properties (absorption), handling, suspending
 - New High reflectivity coatings:
 - Challenges on material selection, doping, deposition procedures, impurities and defects, thermo-mechanical and optical properties, ageing
 - New lasers:
 - Current GW detectors uses 1064nm λ , ET-LF will use 1550nm (or $\sim 2\mu\text{m}$) λ lasers
 - New opto-electronics:
 - Faraday isolators, electro-optical modulators, fibres, photo-sensing

New challenges in Mechanics and Cryogenics

- ET Seismic Filtering will require to push forward 2nd generation technology:
 - Active and passive seismic filtering to have high efficiency, more compact suspension chains
- ET Cryogenics:
 - Challenges on the capability of cooling few hundreds kg test masses
 - without disturbing the interferometer
 - Reducing the cooling time
 - Avoiding the contamination of the mirrors
 - Challenges in the realisation of the payloads compatible both with the cryogenic and operation requirements (geometry, materials,)

New challenges in sensing and actuation

- Current detectors are dominated (at low frequency) by “technical” noises
- ET has science targets in this low-frequency regime
- The low frequency (but not only the LF) is the realm of the noises generated in the sensing-feedback-actuation control loop chain
- ET needs to develop
 - Low noise sensing devices
 - New control methods (Machine Learning?)
 - Low noise actuation tools
 - Thermal compensation systems
 - Adaptive optics

A network of new R&D infrastructures

- The technological evolution requested by ET is stimulating the growth of a series of new facilities and infrastructures where ET R&D is performed:
 - ET pathfinder

Inauguration 8
November 2021

ETpathfinder

Slide: Jan-Simon Henning

- 10m prototype facility, currently under construction in Maastricht
- 14.5M€ investment
- ~20 universities and research institutes from NL/BE/DE/F contribute

A network of new R&D infrastructures

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- ET pathfinder
- Amaldi Centre, Rome

3G Gravitational-Wave Lab

With ARC funds, we are preparing a lab for low temperature tests on a real size prototype of an ET LF-Payload

Cryogenic Tests Area:
 Test Cryostat for a full size LF-Payload, cooled by two PT (~Ø 3 m x 3.5 m):

- 2 thermal shields in insulation vacuum
- 1 experimental chamber with separated vacuum

Pulse Tube Cooling Station

Payload Development and Test Area (LF Payload – Real size)

The Rome1 ET Group:

From Virgo:

| | | |
|-----------|-------------|-----------------------|
| Sibilla | Di Pace | (Post Doc Researcher) |
| Ettore | Majorana | (Full Professor) |
| Valentina | Mangano | (Post Doc Researcher) |
| Luca | Naticchioni | (INFN Researcher) |
| Maurizio | Perciballi | (INFN Technician) |
| Paola | Puppo | (INFN Researcher) |
| Piero | Rapagnani | (Associate Professor) |
| Fulvio | Ricci | (Full Professor) |

From CUORE:

| | | |
|---------|-----------|----------------------------|
| Angelo | Cruciani | (INFN Researcher) |
| Antonio | D'Addabbo | (Post Doc Researcher LNGS) |
| Stefano | Pirro | (INFN Researcher) |

From EGO:

| | |
|-------------|------------------|
| Paolo Ruggi | (EGO Researcher) |
|-------------|------------------|

Grant: About 11M€

A network of new R&D infrastructures

• The technological evolution requested by ET is stimulating the growth of a series of new facilities and infrastructures where ET R&D is performed:

- ET pathfinder
- Amaldi Centre, Rome
- Sar-Grav lab in Sardinia
- Other sites under preparation



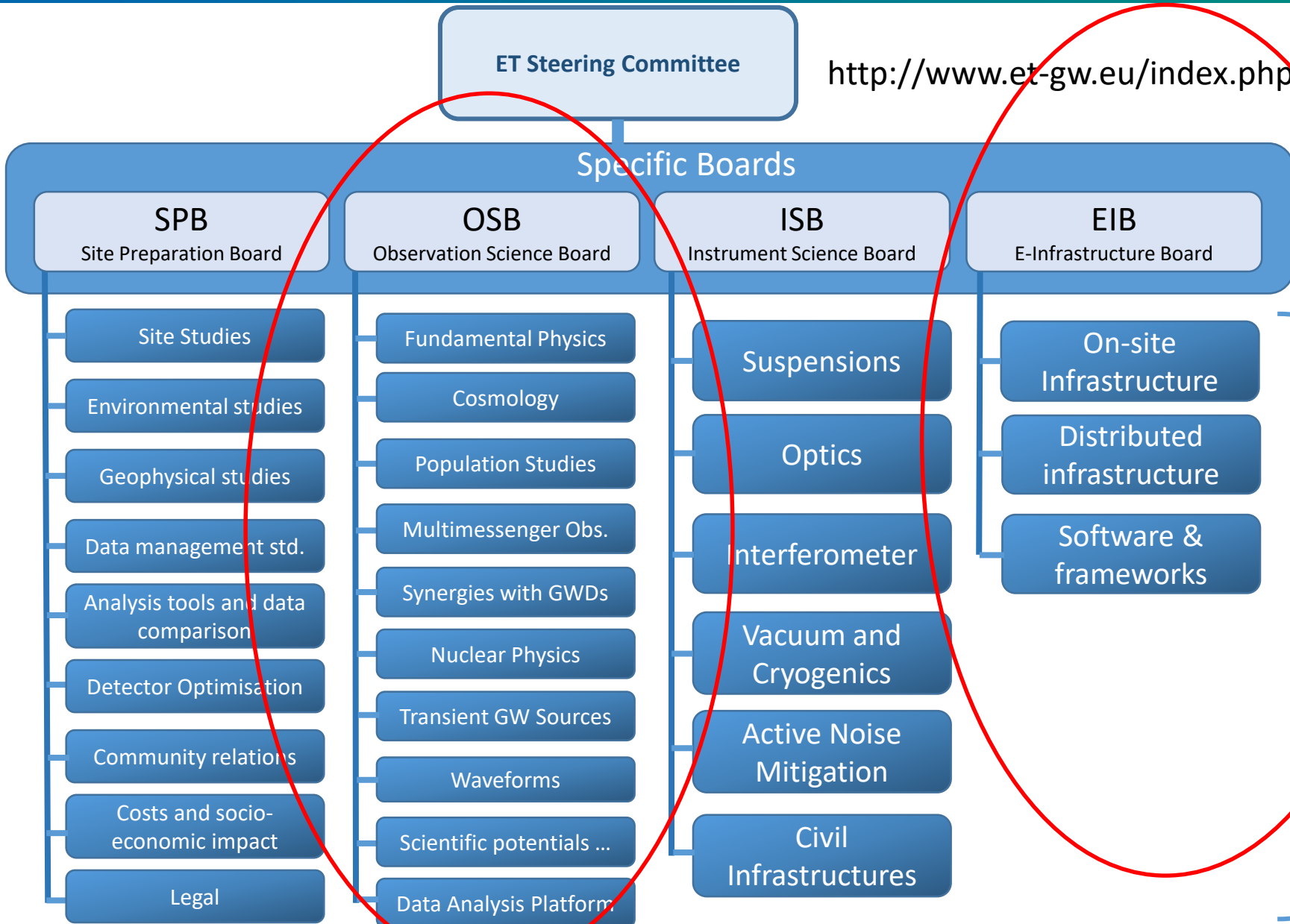
The SarGrav Laboratory

Founded with 3.5 M€ by the Regione Autonoma della Sardegna (RAS) to host low seismic noise underground experiments (low seismic noise experiments, cryogenic payloads, low frequency and cryogenic sensor development)

- ~ 900 m² surface Laboratory
- 3 Underground stations equipped for measurements at different depths
- ~ 50 m² underground area available
- planned a 250 m² underground Lab
- First experiment: Archimedes (founded by INFN)



ET collaboration current organisation



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The proto-collaboration is currently the most organised component of the ET project

Divisions

DAQ, Data stream conditioning, Computing

- GW data is a continuous stream output where the signal is embedded in the noise
 - Challenges: Data Acquisition technologies, data filtering and conditioning
- Low latency analysis and GW alerts
 - Crucial technologies for the Multi-Messenger-Astronomy
 - Efficient computing
 - GPUs, HPC and parallel algorithms
 - Machine Learning and Artificial Intelligence
 - Data Distribution
- Data preservation
 - Technologies and methods
 - Virtual Observatory

Summarizing

1) Photo-sensing, optics, electronics, opto-electronics, quantum technologies

2) Monitoring, sensing and actuation:

- Monitoring and/or sensing
- Control and actuation
- Adaptive systems and compensation

- ET
- CTA
- SKA
- KM3NeT
-

3) Surface and material technologies, coatings

4) Efficient computing and algorithms:

- ML and AI
- High efficiency and low carbon solutions (GPUs, FPGA, ..)
- Low latency alerts
- Data distribution
- Data preservation

Proposed approach

- The richness of the Astro-particle research, under the APPEC hat, is given by the diversity and heterogeneity of technologies to capture the “sidereus nuncii”
- The proposal should match this peculiarity of our research field with the EU requests:
 - To have a part of the proposal aiming to build a common technological platform between the different ESFRI projects
 - Keywords: integration, synergies
 - To devote a fraction of the proposal to develop project specific technologies
 - Keywords: specialisation, cross-fertilisation
 - Balancing to be defined

Proposed roadmap

- 27/10 – 03/11 multilateral discussions
- ~4/11 first meeting of the XXXX proposal
 - Agreement on the topics and first definition of the WPs
 - Definition of the involved ESFRI RI
 - Definition of the core beneficiaries
- ~16/11 Second meeting of the XXXX proposal
 - Definition of the management team
 - Definition of the WPs list
 - Draft definition of the writing team
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